

Physics 211 - Problem Set #5

Solutions

1.) a.) For an interface in 2-dimensions



the length of the interface is
$$L = \int [(dx)^2 + (dz)^2]^{1/2}$$
$$= \int dx \left[1 + \left(\frac{dh}{dx} \right)^2 \right]^{1/2}$$

For an interface $z = h(x)$ in 3-dimensions, the area of the interface is this value times the extent in y .

Then the interfacial tension is

$$\Delta E = \gamma \int dx dy \left[1 + \left(\frac{dh}{dx} \right)^2 \right]^{1/2}$$
$$\approx \gamma \int dx dy \left[1 + \frac{1}{2} \left(\frac{dh}{dx} \right)^2 + \dots \right]$$

For a general 2-d surface in 3 dimensions, a similar formula holds. The derivation is harder; we must divide the surface into parallelograms and compute their areas.

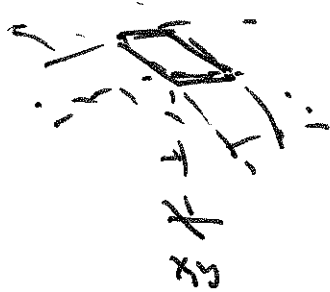
There is an "automatic" method to do this. Write the path length on the interface as

$$\begin{aligned} ds^2 &= g_{ij} dx^i dx^j \\ &= \left| \hat{x} dx + \hat{y} dy + \hat{z} \left(\frac{\partial h}{\partial x} dx + \frac{\partial h}{\partial y} dy \right) \right|^2 \\ &= \left[1 + \left(\frac{\partial h}{\partial x} \right)^2 \right] dx^2 + 2 \frac{\partial h}{\partial x} \frac{\partial h}{\partial y} dx dy + \left[1 + \left(\frac{\partial h}{\partial y} \right)^2 \right] dy^2 \end{aligned}$$

so that

$$g_{ij} = \begin{pmatrix} 1 + \left(\frac{\partial h}{\partial x} \right)^2 & \frac{\partial h}{\partial x} \frac{\partial h}{\partial y} \\ \frac{\partial h}{\partial x} \frac{\partial h}{\partial y} & 1 + \left(\frac{\partial h}{\partial y} \right)^2 \end{pmatrix}$$

If we diagonalize g , it has eigenvalues λ_1^2, λ_2^2 at x, y . A small parallelogram at x, y has area $\lambda_1 \lambda_2 dx dy$



This is $d(\text{area}) = \sqrt{\det g} dx dy$. We can compute $\det g$ and expand it to leading order in h .

$$\begin{aligned} \sqrt{\det g} &= \left[\left(1 + \left(\frac{\partial h}{\partial x} \right)^2 \right) \left(1 + \left(\frac{\partial h}{\partial y} \right)^2 \right) - \left(\frac{\partial h}{\partial x} \frac{\partial h}{\partial y} \right)^2 \right]^{1/2} \\ &\cong 1 + \frac{1}{2} \left(\frac{\partial h}{\partial x} \right)^2 + \frac{1}{2} \left(\frac{\partial h}{\partial y} \right)^2 + \dots \end{aligned}$$

More formally if $g = 1 + \Delta$

$$[\det g]^{\frac{1}{2}} = [1 + \text{tr } \Delta + \dots]^{\frac{1}{2}} = 1 + \frac{1}{2} \text{tr } \Delta + \dots$$

Either way

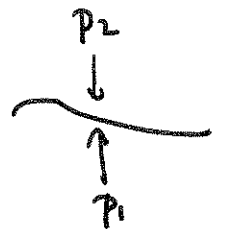
$$\Delta E = \gamma \int dx dy [1 + \frac{1}{2} (\nabla_{\perp}^2 h)^2 + \dots]$$

A small variation of the surface changes the energy by

$$\begin{aligned} \delta E &= \gamma \int dx dy \nabla_{\perp} h \cdot \nabla \delta h \\ &= \int dx dy (-\gamma \nabla_{\perp}^2 h) \delta h \end{aligned}$$

A small variation of the surface changes the energy by this effect and by doing work against the pressure difference

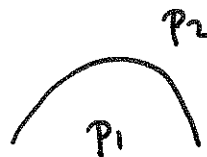
$$\delta E = \int dx dy \delta h [(p_2 - p_1) - \gamma \nabla_{\perp}^2 h]$$



so the equilibrium is

$$p_2 - p_1 - \gamma \nabla_{\perp}^2 h = 0$$

A surface bent like:



has $\frac{d^2 h}{dx^2} < 0$ and so is in equilibrium with a $p_1 > p_2$.

b.) In our analysis of the Kelvin-Helmholtz instability, we had three boundary conditions:

$$\frac{\partial}{\partial z} \delta\phi_2 = \frac{\partial \eta}{\partial t} + V_2 \frac{\partial \eta}{\partial x}$$

$$\frac{\partial}{\partial z} \delta\phi_1 = \frac{\partial \eta}{\partial t} + V_1 \frac{\partial \eta}{\partial x}$$

$$\rho_1 \left[V_1 \frac{\partial}{\partial x} \delta\phi_1 + \frac{\partial}{\partial t} \delta\phi_1 + g\eta \right] \\ = \rho_2 \left[V_2 \frac{\partial}{\partial x} \delta\phi_2 + \frac{\partial}{\partial t} \delta\phi_2 + g\eta \right]$$

the last condition was $-p_1 = -p_2$, so the surface tension modifies it to

$$\rho_1 \left[\left(\frac{\partial}{\partial t} + V_1 \frac{\partial}{\partial x} \right) \delta\phi_1 + g\eta \right] = \rho_2 \left[\left(\frac{\partial}{\partial t} + V_2 \frac{\partial}{\partial x} \right) \delta\phi_2 + g\eta \right] \\ + \gamma \nabla^2 \eta$$

let

$$\delta\phi_1 = \phi_1 e^{-i\omega t + ikx + kz}$$

$$\delta\phi_2 = \phi_2 e^{-i\omega t + ikx - kz}$$

$$\eta = \eta e^{-i\omega t + ikx}$$

as before:

$$\phi_2 = +i \left(\frac{\omega - V_2 k}{k} \right) \zeta \quad \phi_1 = -i \left(\frac{\omega - V_1 k}{k} \right) \zeta$$

so

$$\left[\rho_2 \left(\frac{\omega - V_2 k}{k} \right)^2 + \rho_1 \left(\frac{\omega - V_1 k}{k} \right)^2 - \gamma k^2 + (\rho_2 - \rho_1) g \right] \zeta = 0$$

The equation can be solved for $\omega(k)$.

$$\text{For } \rho_1 = \rho_2 = \rho \quad V_1 = V_2 = 0$$

$$2\rho \frac{\omega^2}{k} - \gamma k^2 = 0$$

$$\omega^2 = \frac{\gamma k^3}{2\rho}$$

$$c.) \quad \text{For } \rho_1 > \rho_2, \quad V_1 = 0$$

$$\begin{aligned} (\rho_1 + \rho_2) \omega^2 - 2\rho_2 V_2 \omega k + \rho_2 V_2^2 k^2 \\ = \gamma k^3 + (\rho_1 - \rho_2) k g \end{aligned}$$

$$\begin{aligned} \left(\omega - \frac{\rho_2 V_2 k}{\rho_1 + \rho_2} \right)^2 = \frac{1}{\rho_1 + \rho_2} \left[\gamma k^3 + (\rho_1 - \rho_2) k g - \rho_2 V_2^2 k^2 \right. \\ \left. + \frac{\rho_2^2 (V_2 k)^2}{\rho_1 + \rho_2} \right] \end{aligned}$$

$$\left(\omega - \frac{\rho_2 V_2 k}{\rho_1 + \rho_2} \right)^2 = \frac{1}{\rho_1 + \rho_2} \left[\gamma k^3 + (\rho_1 - \rho_2) k g - \frac{\rho_1 \rho_2}{\rho_1 + \rho_2} V_2^2 k^2 \right] \quad 6$$

$$\omega = \frac{\rho_2 V_2 k}{\rho_1 + \rho_2} \pm \frac{1}{(\rho_1 + \rho_2)^{1/2}} \left[(\rho_1 - \rho_2) k g - \frac{\rho_1 \rho_2 (V_2 k)^2}{\rho_1 + \rho_2} + \gamma k^3 \right]^{1/2}$$

When the quantity in the bracket becomes negative we set

$$\omega = \frac{\rho_2 V_2 k}{\rho_1 + \rho_2} \pm i \Omega$$

one of which is an unstable mode. For $\gamma = 0$, the bracket is always negative for sufficiently large k .

For $\gamma > 0$, $V_2 = 0$, the bracket is always positive,

In general, the criterion for stability is

$$\frac{\rho_1 \rho_2}{\rho_1 + \rho_2} (V_2 k)^2 < \gamma k^3 + (\rho_1 - \rho_2) k g$$

$$V_2^2 < \left(\frac{\rho_1 + \rho_2}{\rho_1 \rho_2} \right) \left[\gamma k + \frac{(\rho_1 - \rho_2) g}{k} \right]$$

The right-hand side has a minimum when

$$\gamma - \frac{(\rho_1 - \rho_2) g}{k^2} = 0$$

$$k = \left[\frac{(\rho_1 - \rho_2) g}{\gamma} \right]^{1/2}$$

so we have stability for all k if

$$V_2^2 < \frac{\rho_1 + \rho_2}{\rho_1 \rho_2} \cdot 2 \cdot [\gamma g (\rho_1 - \rho_2)]^{1/2}$$

For $\rho_1 \gg \rho_2$

$$V_2 < \left[\frac{2}{\rho_2} \right]^{1/2} [\gamma g \rho_1]^{1/4}$$

d.) For wind-generated waves on the sea, the instability sets in at

$$V_2 = \left[\frac{2}{1.25 \text{ kg/m}^3} \right]^{1/2} \left[(0.074 \text{ kg/sec}^2) (9.8 \text{ m/sec}^2) (1000 \text{ kg/m}^3) \right]^{1/4}$$

(units are: $\frac{[\text{kg}^2/\text{m}^2 \text{ sec}^4]^{1/4}}{[\text{kg}/\text{m}^3]^{1/2}} = \text{m/sec}!$)

$$V_2 = 6.6 \text{ m/sec}$$

$$\lambda_{\text{crit}} = \frac{2\pi}{k_{\text{crit}}} = 2\pi \left[\frac{\gamma}{\rho_1 g} \right]^{1/2} = 2 \text{ mm}$$

stronger winds are needed to generate long-wavelength surface waves.

2.) a.) The volume of the cylinder is $\pi a^2 l$.

Its area is $2\pi a l$. The change in energy if the radius is changed $a \rightarrow a + da$ is

$$dE = - (p - p_0) 2\pi a l da + \gamma 2\pi l da$$

The equilibrium condition is $dE = 0$

$$\text{or } (p - p_0) = \frac{\gamma}{a}$$

b.) From problem 1, the area of the cylinder is given, to quadratic order in small quantities,

by

$$\begin{aligned} \text{Area} &= \int d\phi dz \left[\eta^2 + \left(\frac{\partial \eta}{\partial \phi} \right)^2 \right] \left[1 + \left(\frac{\partial \eta}{\partial z} \right)^2 - \dots \right]^{\frac{1}{2}} \\ &= \int d\phi dz \left(\eta + \frac{1}{2} \frac{1}{\eta} \left(\frac{\partial \eta}{\partial \phi} \right)^2 + \frac{1}{2} \eta \left(\frac{\partial \eta}{\partial z} \right)^2 + \dots \right) \end{aligned}$$

$$\eta = a + s\eta$$

$$= \int d\phi dz a \left[1 + \frac{s\eta}{a} + \frac{1}{2} \frac{1}{a^2} \left(\frac{\partial s\eta}{\partial \phi} \right)^2 + \frac{1}{2} \left(\frac{\partial s\eta}{\partial z} \right)^2 + \dots \right]$$

Now change $s\eta$ by $d\eta$. The change in energy

is

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$$dE = \int d\phi dz \left\{ - (p-p_0) (a+\delta h) dh - \delta p a dh + \gamma \left[dh - dh \frac{1}{a} \frac{\partial^2 \delta h}{\partial \phi^2} - a dh \frac{\partial^2 \delta h}{\partial z^2} \right] \right\}$$

or, for equilibrium:

$$\delta p = - \frac{(p-p_0)(a+\delta h)}{a} + \frac{\gamma}{a} \left[1 - \frac{1}{a} \frac{\partial^2 \delta h}{\partial \phi^2} - a \frac{\partial^2 \delta h}{\partial z^2} \right]$$

$$= - \frac{\gamma}{a} \left(1 + \frac{\delta h}{a} \right) + \frac{\gamma}{a} - \frac{\gamma}{a^2} \frac{\partial^2 \delta h}{\partial \phi^2} - \gamma \frac{\partial^2 \delta h}{\partial z^2}$$

$$\delta p = - \gamma \left[\frac{\delta h}{a^2} + \frac{\partial^2}{\partial z^2} \delta h + \frac{1}{a^2} \frac{\partial^2}{\partial \phi^2} \delta h \right]$$

The boundary condition on the velocity is that the radial velocity at the surface should match the motion of the surface.

So

$$v_r = \frac{\partial}{\partial t} \delta h$$

For $\eta=0$

$$\frac{\partial \vec{v}}{\partial t} + (\vec{v} \cdot \nabla) \vec{v} = - \frac{1}{\rho} \nabla p \quad \nabla \cdot \vec{v} = 0$$

Imaginary:

$$\frac{\partial \vec{v}}{\partial t} = - \frac{1}{\rho} \nabla \delta p \quad \nabla \cdot \vec{v} = 0$$

e.) Assine potential flow $\vec{v} = \nabla\psi$

$$\text{Then } \vec{v} \cdot \vec{v} = 0 \Rightarrow \nabla^2\psi = 0$$

$$\frac{\partial \vec{v}}{\partial t} = -\frac{1}{\rho} \nabla \delta p \Rightarrow \frac{\partial}{\partial t} \nabla\psi = -\frac{1}{\rho} \nabla \delta p$$

$$\nabla \delta p = -\rho \frac{\partial}{\partial t} \nabla\psi = 0$$

Let $\delta p = p(r) e^{-i\omega t + ikz + in\phi}$ etc

$$\nabla^2 \delta p = 0 \Rightarrow \frac{1}{r} \frac{\partial}{\partial r} r \frac{\partial}{\partial r} p(r) - k^2 p(r) - \frac{n^2}{r^2} p(r) = 0$$

again

$$\left[\frac{1}{r} \frac{d}{dr} r \frac{d}{dr} - \left(k^2 + \frac{n^2}{r^2} \right) \right] p(r) = 0$$

d.) This is the equation of the modified Bessel functions $I_n(kr)$, $K_n(kr)$. $p(r)$ must be regular at $r=0$,

hence $p(r) = A I_n(kr)$

Now reconstruct \vec{v}

$$\frac{\partial \vec{v}}{\partial t} = -\frac{\nabla p}{\rho} \quad \text{so} \quad -i\omega \vec{v}(r) = -\frac{1}{\rho} \nabla p(r)$$

For v_r

$$\begin{aligned}v_r &= -i \frac{1}{\omega \rho} \frac{\partial}{\partial r} p(r) \\ &= -i \frac{k}{\omega \rho} A I_n'(kr)\end{aligned}$$

also

$$\begin{aligned}v_r &= \frac{\partial}{\partial t} \zeta_h \quad \text{at } r=a \\ &= -i\omega \zeta_h\end{aligned}$$

$$\text{so } \zeta_h = \frac{k}{\omega^2 \rho} A I_n'(ka)$$

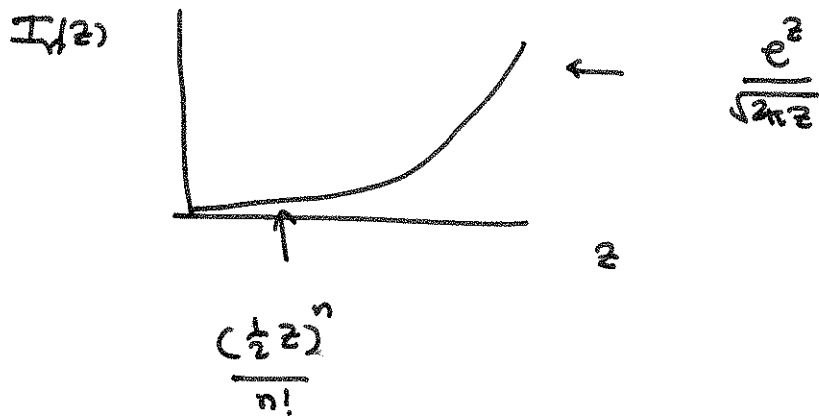
We can now write the boundary condition on the pressure on p. 9
as

$$A I_n(ka) = -\gamma \frac{k}{\omega^2 \rho} A I_n'(ka) \left[\frac{1}{a^2} - k^2 - \frac{n^2}{a^2} \right]$$

so that

$$\omega^2 = \left[(ka)^2 + (n^2 - 1) \right] \frac{\gamma k}{a^2 \rho} \frac{I_n'(ka)}{I_n(ka)}$$

e.) You can look up the behavior of $I_n(z)$ in your favourite handbook. You will find:



$I_n(z) > 0$ $I'_n(z) > 0$ for z real and positive, so

$I'_n/I_n > 0$. Then ω^2 is positive if

$$(ka)^2 + n^2 - 1 > 0$$

there is no instability for $n = \pm 1, \pm 2, \dots$

But, for $n=0$, there is an instability if

$$(ka)^2 - 1 < 0 \quad \text{or} \quad k < \frac{1}{a}$$

3.) This problem is an application of the stability analysis for inviscid Couette flow done in class. We derived, for axisymmetric perturbations,

$$(-DD_+ + k^2) V_r = \frac{k^2}{\omega^2} \Phi V_r$$

where $\Phi = \frac{1}{r^3} \frac{d}{dr} (rV)^2$ $D = \frac{d}{dr}$ $D_+ = \frac{d}{dr} + \frac{1}{r}$

$\rightarrow V_r = 0$ on the boundaries.

For a bulk rotation $V = \Omega r$, $\Phi = 4\Omega^2$

Then the above equation becomes:

$$\left[-\frac{d}{dr} \left(\frac{d}{dr} + \frac{1}{r} \right) + k^2 - k^2 \frac{\Omega^2}{\omega^2} \right] V_r = 0$$

$$\stackrel{\Omega}{\rightarrow} \left[-\frac{d}{dr} \frac{1}{r} \frac{d}{dr} r + k^2 \left(1 - \frac{\Omega^2}{\omega^2} \right) \right] V_r = 0$$

write $V_r = \frac{d}{dr} f$; then f satisfies

$$\left[-\frac{1}{r} \frac{d}{dr} r \frac{d}{dr} + k^2 \left(1 - \frac{\Omega^2}{\omega^2} \right) \right] f = 0$$

This is Bessel's equation. The eigenfunctions are!

$$J_0(xr), \quad x^2 = k^2 \left(\frac{\Omega^2}{\omega^2} - 1 \right)$$

since

$$\left[\frac{1}{r} \frac{d}{dr} r \frac{d}{dr} + x^2 \right] J_0(xr) = 0$$

then

$$\frac{\Omega^2}{\omega^2} = k^2 + x^2$$

$$\omega^2 = \frac{k^2}{k^2 + x^2} \Omega^2$$

x is determined by the condition

$$v_r = 0 \text{ at } r = a \Rightarrow J_0'(xa) = 0$$

[I have also imposed that v_r is regular at $r=0$; that is why J_0 and not Y_0 is used.]

Now $J_0'(z) = -J_1(z)$

so if z_m is the m^{th} zero of $J_1(z)$

$$\omega_m = \frac{k}{[k^2 + z_m^2/a^2]^{1/2}} \Omega$$