

Physics 152 Lecture 2b

2b.1

Relativistic Kinematics

First, a word about units

In particle physics, it is traditional to set

the speed of light $c = 1$

and (reduced) Planck's constant $\hbar \equiv \frac{h}{2\pi} = 1$.

This is the same as converting times and distances into energies, for which the standard unit is

$1 \text{ GeV} = 10^9 \text{ electron volts.}$

Convenient because the rest energy of the proton is $(m_p c^2 = 0.938 \text{ GeV.})$

Times \rightarrow distances using $c \approx 3 \times 10^{10} \text{ cm s}^{-1} = 1 \text{ foot ns}^{-1} = 30 \text{ cm ns}^{-1}$

Example: Several unstable particles (B's, D's, ...) have lifetimes $\tau \approx 1 \text{ ps}$

$\rightarrow c\tau \approx 0.3 \text{ mm}$ \leftarrow (very approximate flight distance before decay, if $v \approx c$)

Microscopic distances:

Atomic physicists often use $1 \text{ \AA} = 10^{-10} \text{ m} = 10^{-8} \text{ cm} \approx r_{\text{atom}}$

Nuclear/particle physicists use $1 \text{ fermi} = 1 \text{ fm} \equiv 1 \text{ femtometer} = 10^{-15} \text{ m} = 10^{-13} \text{ cm} \approx r_{\text{proton}}$

Distances \rightarrow Energies using

$1 = \hbar c = 197 \text{ MeV fm (nuclear)} = 0.197 \text{ GeV fm (particle)}$

$$[L] = [T] = \frac{1}{[E]}$$

"Cross sections" \equiv area \rightarrow probability of interaction measured by nuclear physicists in "barns" $1 \text{ b} = 10^{-24} \text{ cm}^2$

Typical particle physics cross sections: $1 \text{ nb} = 10^{-33} \text{ cm}^2$, $1 \text{ pb} = 10^{-36} \text{ cm}^2$, $1 \text{ fb} = 10^{-39} \text{ cm}^2$
 $1 = (\hbar c)^2 = 3.89 \times 10^8 \text{ GeV}^2 \text{ pb}$ \leftarrow Use to convert $\text{GeV}^2 \rightarrow \text{pb}$

Special Relativity

• Laws of physics apply in any reference frame, provided that we make appropriate transformations of coordinates (and other quantities).

• Speed of light = constant, c in all frames.

(Michelson-Morley experiment)

"Galilean transformations"

$$\begin{cases} t' = t \\ x' = x - vt \end{cases}$$

2 frames with rel. velocity v in x direction

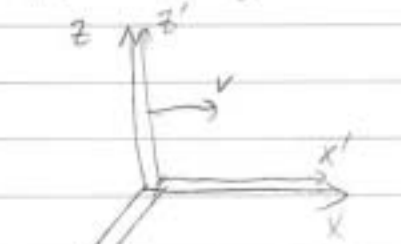
no good

- leads to wrong velocity addition formula, $u^{\oplus} = u' + v$

• For $u = c \Rightarrow u' = c + v \neq c$

• Need to generalize to $t' = at + bx$

$$x' = dt + ex$$

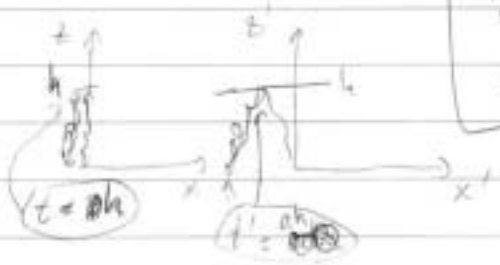


$$\begin{cases} y' = y \\ z' = z \end{cases}$$

Light in $\pm x$ direction
 for $t' = x'$
 $\Rightarrow a + b = d + e$
 for $t' = -x'$
 $\Rightarrow -a + b = d - e$
 $\Rightarrow d = b, e = a$

to avoid paradox: does ball get through cylinder or not?

$\beta \equiv \frac{v}{c}$ often



$$h' = \frac{h}{\gamma} = h \sqrt{1 - \frac{v^2}{c^2}} \Rightarrow \sqrt{1 - \frac{v^2}{c^2}} = \frac{h'}{h}$$

Satisfied by $a = \gamma \equiv \frac{1}{\sqrt{1 - \frac{v^2}{c^2}}}$, $b = -v\gamma$

$$\sqrt{1 + \frac{v^2}{c^2}} = \sqrt{1 + \frac{v^2}{c^2}} = \frac{1}{\sqrt{1 - \frac{v^2}{c^2}}}$$

$$\Rightarrow \begin{cases} t' = \gamma(t - vx/c^2) \\ x' = \gamma(-vt + x) \\ y' = y \\ z' = z \end{cases}$$

[reduces to Galilean for $v \ll c$]

Inverse transformation $\Leftrightarrow v \leftrightarrow -v$

2b.3

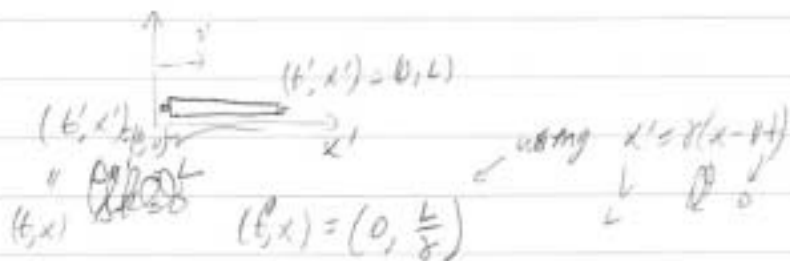
Consequences:

$$\Delta t' = \gamma(\Delta t - v \Delta x)$$

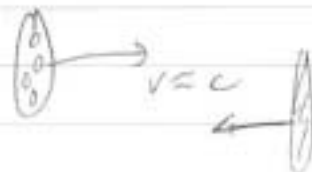
(1) so $\Delta t = 0$ can be $\Delta t' \neq 0$ "relativity of simultaneity"

(2) Lorentz contraction

moving stick,
length L in rest frame
looks like length L/γ .



\Rightarrow Protons & nuclei "pancake"
at high energies.



(3) Time dilation.

$$(t', x') = (0, 0) \rightarrow (t', x') = (\tau, 0)$$

$$(t, x)$$

$$\Delta t = \gamma(\Delta t' + vx')$$

$$\Rightarrow \Delta t = \gamma \tau$$

Particle lifetimes
are longer by γ .

Average flight length is $\Delta x = \beta \Delta t = \beta \gamma c \tau$

(can be $\gg c \tau$ for $\beta \approx 1$)

(e.g. B mesons with $c \tau \approx 0.3 \text{ mm}$
can fly several nm before decaying.)

Velocity addition formula becomes

$$u = \frac{u' + v}{1 + uv/c^2}$$

$$u = c = \beta \Rightarrow u = \frac{1+v}{1+v} = 1 \quad \checkmark$$

4-vectors

$x^0 = t, x^1 = x, x^2 = y, x^3 = z$

Write Lorentz transformation as

x^μ

$x^{\mu'} = \Lambda^{\mu'}_{\nu} x^\nu$

implicit $\sum_{\nu=0}^3$

where for v in x -direction,

$\Lambda^{\mu'}_{\nu} = \begin{pmatrix} \gamma & -\gamma v & 0 & 0 \\ -\gamma v & \gamma & 0 & 0 \\ 0 & 0 & 1 & 0 \\ 0 & 0 & 0 & 1 \end{pmatrix}$

Anything transforming like x^μ in \otimes is called a covariant 4-vector.

Recall proper time interval $\Delta\tau$ time elapsed on clock travelling along from $(0,0,0,0) \rightarrow (\Delta x^0, \Delta x^1, \Delta x^2, \Delta x^3)$

$(\Delta\tau)^2 = (\Delta x^0)^2 - (\Delta x^1)^2 - (\Delta x^2)^2 - (\Delta x^3)^2 = \eta_{\mu\nu} \Delta x^\mu \Delta x^\nu$

where $\eta_{\mu\nu} = \begin{pmatrix} 1 & 0 & 0 & 0 \\ 0 & -1 & 0 & 0 \\ 0 & 0 & -1 & 0 \\ 0 & 0 & 0 & -1 \end{pmatrix}$ is Minkowski metric

Used to raise/lower indices; e.g. $X_\mu \equiv \eta_{\mu\nu} X^\nu$ (called a covariant 4-vector)

$\eta_{\mu\nu}$ is an invariant tensor,

$\Lambda^{\mu'}_{\alpha} \Lambda^{\nu'}_{\beta} \eta_{\mu\nu} = \begin{pmatrix} \gamma & -\gamma v \\ -\gamma v & \gamma \end{pmatrix} \begin{pmatrix} 1 & 0 \\ 0 & -1 \end{pmatrix} \begin{pmatrix} \gamma & -\gamma v \\ -\gamma v & \gamma \end{pmatrix} = \begin{pmatrix} \gamma & -\gamma v \\ -\gamma v & \gamma \end{pmatrix} \begin{pmatrix} \gamma & -\gamma v \\ -\gamma v & \gamma \end{pmatrix} = \begin{pmatrix} 1 & 0 \\ 0 & -1 \end{pmatrix} = \eta_{\alpha\beta}$

True for all Λ , in fact it defines

the group of Lorentz transformations: $SO(1,3)$ (real matrices) preserves (+---)

(2b.5)

4-vectors can be combined into Lorentz invariant quantities (Lorentz scalars) using $\eta_{\mu\nu}$ (or $\eta^{\mu\nu} = \eta_{\mu\nu}^{-1} = \eta_{\mu\nu}$)

$$a \cdot b \equiv a^\mu b_\mu = \eta_{\mu\nu} a^\mu b^\nu$$

Lorentz \rightarrow

$$\eta_{\mu\nu} \Lambda^\mu_\alpha \Lambda^\nu_\beta a^\alpha b^\beta = \eta_{\alpha\beta} a^\alpha b^\beta = a \cdot b$$

$$a \cdot b = a^0 b^0 - \vec{a} \cdot \vec{b}$$

$$a^2 = (a^0)^2 - \vec{a}^2 \begin{cases} > 0 & \text{timelike} \\ = 0 & \text{lightlike} \\ < 0 & \text{spacelike} \end{cases}$$

• Tensors transform with more indices

$$T^{\mu\nu} \rightarrow \Lambda^\mu_\alpha \Lambda^\nu_\beta T^{\alpha\beta}, \dots$$

Energy & Momentum of particles

• Ordinary velocity $\vec{v} \equiv \frac{d\vec{x}}{dt} = \frac{d\vec{x}}{dx^0}$

does not have

simple Lorentz transformation properties because of x^0 in denominator.

Better: to put proper time interval $d\tau$ in denominator. ($dt = \gamma d\tau$) (time dilation)

$\Rightarrow \eta^i \equiv \frac{dx^i}{d\tau} = \gamma \frac{dx^i}{dt}$ will transform like 3 of the 4-components of a 4-vector.

The 4th component is

$$\eta^0 \equiv \frac{dx^0}{d\tau} = \frac{dt}{d\tau} = \gamma$$

\Rightarrow 4-velocity $\eta^\mu \equiv \gamma(1, v_x, v_y, v_z)$

Lorentz invariant magnitude of 4-velocity: $\eta^\mu \eta_\mu = \gamma^2(1 - v^2) = 1$

Momentum of a particle

- Because $\vec{v} = \frac{d\vec{x}}{dt}$ does not have simple transformation properties, neither does $\vec{p} = m\vec{v}$.
- Better to use $\vec{p} = m\vec{\eta} = m \frac{d\vec{x}}{d\tau} = m\gamma\vec{v}$.
- Agrees with old \vec{p} for $v \ll 1$.

- The full 4-vector is $\textcircled{*} p^\mu = m\eta^\mu = m\gamma(1, v_x, v_y, v_z)$
 - The time component $p^0 \equiv E = m\gamma$ ← (total energy of particle: rest energy + kinetic energy)
 $(E_{\text{rest}} = mc^2)$
- $p^\mu = (E, p_x, p_y, p_z)$

$p^\mu p_\mu$ is Lorentz invariant. Evaluate in particle rest frame $\Rightarrow p^\mu = (M, 0, 0, 0)$
 or directly from $(*)$. (rest energy only)

$\Rightarrow p^\mu p_\mu = p^2 = p^\mu p_\mu = M^2$ ← "on-shell condition" (mass)

• Thus $E^2 - \vec{p}^2 = M^2$

• Can also expand for small $v \Rightarrow E = m\gamma = \frac{m}{\sqrt{1-v^2}} = m + \frac{1}{2}mv^2 + \frac{3}{8}mv^4 + \dots$
 rest energy (nonrelativistic kinetic energy) (relativistic correction)

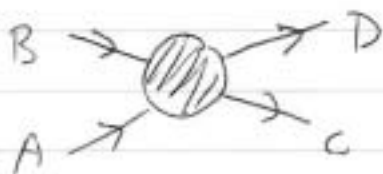
• As $v \rightarrow c$, m becomes inconvenient and irrelevant. $\gamma \rightarrow \infty$
 $\vec{p} = m\vec{v}\gamma \gg m$
 $E = m\gamma \gg m$.

$E^2 - \vec{p}^2 \approx 0$, or $E \approx |\vec{p}|$ ← massless (γ) or high energy particle

• Difference between 2 eV and 3 eV photon is not the mass, but (via quantum mechanics) the frequency, $E = h\nu$, or wavelength, $E = hc/\lambda$

Collisions

$A+B \rightarrow C+D$ ~~elastic~~ "2 \rightarrow 2" (most frequent)



"free particles at infinity" $\Rightarrow p_i = \text{constants at } \infty$
 $i = A, B, C, D$

- Collision takes place over a very short time/distance interval \Rightarrow ignore all external forces & fields.
- Then energy-momentum conservation

$$\Rightarrow \boxed{p_A^{\mu} + p_B^{\mu} = p_C^{\mu} + p_D^{\mu}} \quad (1)$$

0th component: $\boxed{E_A + E_B = E_C + E_D}$ (2)

total energy is conserved.

- But this includes rest energy, which can change if $m_A + m_B \neq m_C + m_D$.
- So total kinetic energy is not conserved.

An extreme example: $e^+e^- \rightarrow \tau^+\tau^-$ at threshold.

with $E_{e^+} = E_{e^-} = 1.8 \text{ GeV}$ $m_{\tau^\pm} = 1.777 \text{ GeV}$

(Eq. (2) + Symmetry) $\Rightarrow E_{\tau^+} = E_{\tau^-} = 1.8 \text{ GeV} = \gamma_\tau \cdot 1.777 \text{ GeV}$

$$\Rightarrow \gamma_\tau = 1.01 = \frac{1}{\sqrt{1-\beta_\tau^2}} \Rightarrow \beta_\tau = \sqrt{1 - \frac{1}{1.01^2}} = 0.16c$$

τ 's are nonrelativistic,

$$\boxed{(K.E.)_\tau = \frac{1}{2} m_\tau v_\tau^2 = 22 \text{ MeV} \ll m_\tau c^2}$$

while

$$(K.E.)_e \approx E_e = 1.8 \text{ GeV}$$

• How many ^{kinematic} quantities characterize the "process" or "reaction" $A + B \rightarrow C + D$?

~~16~~

• Not counting m_A, \dots, m_D which are attributes of the particles, not the process.

• Start with $4 \times 4 = 16$ quantities,
 $p_A^x \quad p_B^x \quad p_C^x \quad p_D^x$

• But $p_i^2 = m_i^2$ removes 4 ($i = A, B, C, D$)
 $p_A^x + p_B^x = p_C^x + p_D^x$ removes 4 more
 $16 - 4 - 4 = 8$.

• Still not done. There are 3 boosts (along x, y, z axes) and 3 rotations (around x, y, z axes)

~~8~~ which only change the frame, not the essential attributes of the reaction.

$8 - 3 - 3 = \boxed{2}$ ← true number of kinematic quantities

• What are they?

• Let's use boost invariance to ~~go~~ go to center of mass frame

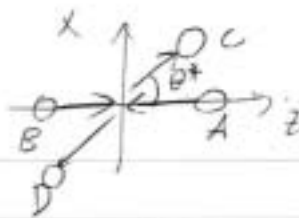
$\vec{p}_{TOT} = \vec{p}_A + \vec{p}_B = 0$
($= \vec{p}_C + \vec{p}_D$)



• Then use rotation invariance to line \vec{p}_A & \vec{p}_B up (the "beamline") along z direction.



• Rotate around z axis until C & D are in (say) xz plane



(26.9)

$\theta^* \equiv \theta_{CM}$ is center of mass scattering angle.

The other quantity is ~~the~~ just the center-of-mass energy, E_{CM}

• Let's count a second way (Mandelstam) construct Lorentz invariants $(p_i \pm p_j)^2$.

problem 3.22 of Griffiths

• How many are there?

~~(4)~~ $i=j$ gives 0 or ~~also~~ $(2p_i)^2 = 4m_i^2$ (we don't count this)

• So take $i \neq j$.

Since $(p_i + p_j)^2 + (p_i - p_j)^2 = 2(p_i^2 + p_j^2) = 2(m_i^2 + m_j^2)$, we can take + or - but not both.

Mandelstam chose

$$\begin{aligned} s &\equiv (p_A + p_B)^2 = (p_C + p_D)^2 \\ t &\equiv (p_A - p_C)^2 = (p_B - p_D)^2 \\ u &\equiv (p_A - p_D)^2 = (p_B - p_C)^2 \end{aligned}$$

with one constraint eqn:

$$0 = (p_A + p_B - p_C - p_D)^2$$

$$= 0 \left(\sum_{i=1}^4 m_i^2 \right) + 2(p_A \cdot p_B + p_C \cdot p_D) - 2(p_A \cdot p_C + p_B \cdot p_D) - 2(p_A \cdot p_D + p_B \cdot p_C)$$

$$= 0 \left(\sum_{i=1}^4 m_i^2 \right) + 2 \left(s - \sum_{i=1}^4 m_i^2 \right) + 2 \left(t - \sum_{i=1}^4 m_i^2 \right) + 2u - \sum_{i=1}^4 m_i^2$$

$$\Rightarrow \boxed{s + t + u = \sum_{i=1}^4 m_i^2}$$

Thus u (say) can be eliminated.

\Rightarrow 2 variables: $\boxed{s, t}$

Now, in CM frame $S = (\underbrace{E_A + E_B}_{E_{CM}}, \vec{0})^2$
 $\Rightarrow S = E_{CM}^2$

t must be a function of E_{CM} and θ_{CM}
 $t = f(E_{CM}, \theta_{CM})$

• In the ~~massive~~ case where ~~the~~
~~particles~~ all particles are massless,
 this relation is particularly simple:

$E = \frac{E_{CM}}{2}$

$p_A = E(1, 0, 0, 1)$
 $p_B = E(1, 0, 0, -1)$
 $p_C = E(1, \sin\theta\cos\phi, \sin\theta\sin\phi, \cos\theta)$
 $p_D = E(1, -\sin\theta\cos\phi, -\sin\theta\sin\phi, -\cos\theta)$
 (t x y z)

~~$p_A = (E, 0, 0, E)$~~
 ~~$p_B = (E, 0, 0, -E)$~~
 ~~$p_C = E(1, \sin\theta, \cos\theta)$~~
 ~~$p_D = E(1, -\sin\theta, -\cos\theta)$~~

$\Rightarrow t = (p_A - p_C)^2 = -2p_A \cdot p_C = -2E^2(1 - \cos\theta)$
 $= -2\left(\frac{E_{CM}}{2}\right)^2(1 - \cos\theta)$

$\Rightarrow t = -\frac{1}{2} E_{CM}^2 (1 - \cos\theta) = -\frac{1}{2} S (1 - \cos\theta)$

$u = -\frac{1}{2} S (1 + \cos\theta)$

2 Body Decay is Simpler

Ex. 3.3 of Griffiths, π at rest, $\pi^+ \rightarrow \mu^+ \nu_\mu$

$m_\mu = 0$. What is velocity of muon? (Also momentum)

Use 4-vectors, $p_\nu^2 = 0$

\Rightarrow rewrite $p_\pi = p_\mu + p_\nu$ as $p_\nu = p_\pi - p_\mu$, square it

$$\Rightarrow 0 = p_\nu^2 = (p_\pi - p_\mu)^2 = m_\pi^2 - 2p_\pi \cdot p_\mu + m_\mu^2$$

$$p_\pi = (m_\pi, \vec{0}) \Rightarrow (p_\pi \cdot p_\mu = m_\pi E_\mu)$$

$$p_\mu = (E_\mu, \vec{p}_\mu)$$

$$\Rightarrow E_\mu = \frac{m_\pi^2 + m_\mu^2}{2m_\pi}$$

$$p_\mu = \sqrt{E_\mu^2 - m_\mu^2} = \sqrt{\frac{(m_\pi^2 + m_\mu^2)^2 - 4m_\pi^2 m_\mu^2}{4m_\pi^2}}$$

$$\Rightarrow p_\mu = \frac{m_\pi^2 - m_\mu^2}{2m_\pi}$$
 ← (measure with \vec{B} field in detector)

$$v_\mu = \frac{p_\mu}{E_\mu} = \frac{m_\pi^2 - m_\mu^2}{m_\pi^2 + m_\mu^2}$$

$m_\pi = 140 \text{ MeV}$
 $m_\mu = 106 \text{ MeV}$

$$v_\mu = 0.271 \leftarrow \text{(semi-relativistic)}$$

• Note that v_μ, p_μ, E_μ are all fixed by the particle masses in a 2-body decay. "Mono-energetic muon".

• In a 3-body decay, like $^{60}\text{Co} \rightarrow ^{60}\text{Ni} + e^- + \bar{\nu}_e$ this is not true: electron can have a spectrum of different energies. Key to Pauli's proposal of the neutrino in 1930.

Use of CM frame

Ex. 3.4 of Griffiths

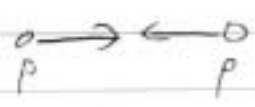
6 Bevatron at Berkeley built to produce antiprotons.

(beam) $p + p \rightarrow p + p + p + \bar{p}$
What is minimum beam energy E_{min} required?

Lab frame



CM frame



- Minimum beam energy \leftrightarrow minimum CM energy
 \leftrightarrow all particles at rest in CM (no wasted K.E.)

$\Rightarrow E_{cm}^{min} = 4m_p$

Now we compute $s = E_{cm}^2$ before collision, in lab frame.

$s = ((E, \vec{p}) + (m, \vec{0}))^2 = (E+m, \vec{p})^2 = (E+m)^2 - p^2$
 $= E^2 + 2m_p E + m_p^2 - (E^2 - m_p^2) = 2m_p(E+m_p)$

So, we have $2m_p(E^{min} + m_p) = 16m_p^2$

$\Rightarrow E^{min} = 7m_p$

• For $E \gg m$, $s \approx 2 E_{beam} m_{target}$ for fixed target expt.



• In colliding beams, ~~$s = 4E^2$~~

$s = 4E_{beam}^2 \gg 2E_{beam} m_{target}$



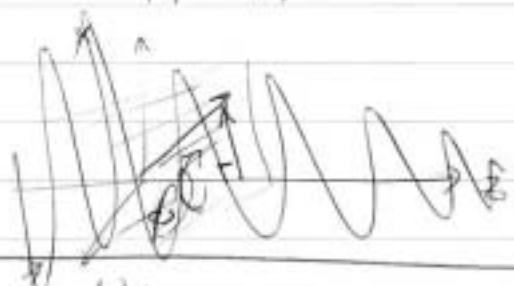
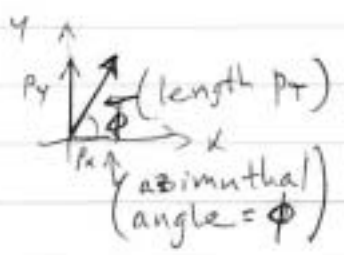
eg. Tevatron 1 TeV + 1 TeV \gg 2000 TeV fixed target !!!

Remark:

Transverse momenta ^{with respect to z} are boost-invariant for a boost along z:

$P_T \equiv \sqrt{p_x^2 + p_y^2}$

$E' = \gamma(E - \beta p_z)$
 $p_z' = \gamma(-\beta E + p_z)$
 $p_x' = p_x$
 $p_y' = p_y$



A better longitudinal variable is sometimes the rapidity

$y = \frac{1}{2} \ln \left(\frac{E + p_z}{E - p_z} \right)$

under a boost

$y \rightarrow y' = \frac{1}{2} \ln \left(\frac{E' + p_z'}{E' - p_z'} \right) = \frac{1}{2} \ln \left(\frac{\gamma(1-\beta)(E + p_z)}{\gamma(1+\beta)(E - p_z)} \right)$

$y' = y + \frac{1}{2} \ln \left(\frac{1-\beta}{1+\beta} \right)$

(p_T, η, ϕ) Used often in hadron colliders, where you don't know the CM frame of the quark collisions

