

Lecture 12 QED finale

12.1 WMA

Amplitudes with external photons

Incoming f^0 ϵ^μ
 Outgoing f^1 $\epsilon^{\mu'}$

Lorentz gauge $\partial_\mu A^\mu = 0 \Rightarrow \epsilon \cdot k = 0$

Still free to shift $(\epsilon^\mu \rightarrow \epsilon^\mu + \text{constant} \cdot k^\mu)$

and get same answer - "Ward identity"
 - consequence of gauge symmetry.

We discussed Coulomb gauge ϵ 's before, $\epsilon \begin{cases} \epsilon^0 = 0 \\ \vec{\epsilon} \cdot \vec{k} = 0 \end{cases}$

E.g. $\epsilon_{(1)}^\mu = (0, \hat{E}_x) = (0, 1, 0, 0)$
 $\epsilon_{(2)}^\mu = (0, \hat{E}_y) = (0, 0, 1, 0)$ for $\vec{k} = |\vec{k}|\hat{z}$
 $k = |\vec{k}|(0, 0, 0, 1)$

$$\Rightarrow \sum_{\text{pol's}} \epsilon_\lambda^\mu \epsilon_\lambda^{\nu*} = \begin{pmatrix} 0 & 0 & 0 & 0 \\ 0 & 1 & 0 & 0 \\ 0 & 0 & 1 & 0 \\ 0 & 0 & 0 & 0 \end{pmatrix} = \delta^{ij} - \hat{k}^i \hat{k}^j$$

~~But if we let $\epsilon_\lambda^\mu = \frac{1}{|\vec{k}|} \epsilon_{\lambda\alpha}^\mu k^\alpha$~~

But this is messy, not covariant.

Let $M(\epsilon^\mu) = \epsilon^\mu M_\mu = (\epsilon^\mu + \text{const.} k^\mu) M_\mu$

$$\Rightarrow k_\mu M_\mu = 0. \text{ Here, } k^\mu M_\mu = |\vec{k}|(M_0^2 - M^2)$$

$$\Rightarrow \mu^0 = M^2$$

$$\Rightarrow \sum_{\text{pol's}} \epsilon_\lambda^\mu \epsilon_\lambda^{\nu*} M^\mu M^{\nu*} = M_0^2 + M^2$$

$$= -\mu_0^2 + M_0^2 + M^2 + M^2$$

$$= -g_{\mu\nu} M^\mu M^{\nu*}$$

Thus $\sum_{\text{pol's}} \epsilon_\lambda^\mu \epsilon_\lambda^{\nu*} = -g_{\mu\nu}$ can always be used for the photon polarization sums.

Example: $e^+e^- \rightarrow \gamma\gamma$

$m_e \rightarrow 0$



$$M = \frac{e^2}{(p_1 - p_3)^2} \bar{v}(2) \not{\epsilon}_4 (\not{p}_1 - \not{p}_3) \not{\epsilon}_3 u(1) + \frac{e^2}{(p_1 - p_4)^2} \bar{v}(2) \not{\epsilon}_3 (\not{p}_1 - \not{p}_4) \not{\epsilon}_4 u(1)$$

First check Ward identity: \odot Let $\underline{\epsilon}_3 = p_3$
result should vanish.

$$M \rightarrow \frac{e^2}{-2p_1 \cdot p_3} \bar{v}(2) \not{\epsilon}_4 (\not{p}_1 - \not{p}_3) \not{p}_3 u(1) + \frac{e^2}{-2p_2 \cdot p_3} \bar{v}(2) \not{p}_3 (-\not{p}_2 + \not{p}_3) \not{\epsilon}_4 u(1)$$

$$= \frac{1-e^2}{-2p_1 \cdot p_3} \bar{v}(2) \not{\epsilon}_4 u(1) (2p_1 \cdot p_3) + \frac{e^2}{-2p_2 \cdot p_3} \bar{v}(2) \not{\epsilon}_4 u(1) (-2p_2 \cdot p_3)$$

$$= 0 \quad \checkmark$$

Casimir's trick + ϵ pol. sum \Rightarrow

$$\langle |M|^2 \rangle = \frac{e^4}{4} \left\{ \frac{1}{(2p_1 \cdot p_3)^2} \text{tr} [\not{p}_2 \not{\epsilon}_4 (\not{p}_1 - \not{p}_3) \not{\epsilon}_3 \not{p}_1 \not{\epsilon}_3 (\not{p}_1 - \not{p}_3) \not{\epsilon}_4] + \frac{1}{(2p_1 \cdot p_3)(2p_2 \cdot p_3)} \text{tr} [\not{p}_2 \not{\epsilon}_4 (\not{p}_1 - \not{p}_3) \not{\epsilon}_3 \not{p}_1 \not{\epsilon}_4 (\not{p}_1 - \not{p}_4) \not{\epsilon}_3] + (3 \leftrightarrow 4) \right\}$$

First trace is $4 \text{tr} [\not{p}_2 (\not{p}_1 - \not{p}_3) \not{p}_1 (\not{p}_1 - \not{p}_3)] = 8 p_1 \cdot p_3 + \text{tr} [\not{p}_2 \not{p}_3]$

Second trace is $\frac{32 p_1 \cdot p_3 p_2 \cdot p_3}{4} = 8 p_1 \cdot p_3 p_2 \cdot p_3 = 8 p_1 \cdot p_3 p_2 \cdot p_3$

$\Rightarrow \langle |M|^2 \rangle = 2 e^4 \left[\frac{p_2 \cdot p_3}{p_1 \cdot p_3} + \frac{1}{4} \right]$

R = +
L = -

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Second example:

$$e^+e^- \rightarrow \gamma\gamma$$

- A clean test of QED at very high energies. Look for new $e^+e^- \gamma\gamma$ interactions.

2 helicity amplitudes: $e_L^+e_R^- \rightarrow \gamma_R\gamma_R$

0 in lowest-order QED for $m_e=0$ it turns out!

$$e_L^+e_R^- \rightarrow \gamma_L\gamma_R$$

$$\mathcal{M}(e_L^+e_R^- \rightarrow \gamma_L\gamma_R) = 2e^2 \sqrt{\frac{u}{t}}$$

$$J_2 = 1 \rightarrow J_2 = \pm 2$$

\Rightarrow more helicity suppression one way than other. \checkmark

$$\Rightarrow \left(\frac{d\sigma}{d\Omega} \right) \propto \frac{u}{t} + \frac{t}{u} = \frac{1+\cos\theta}{1-\cos\theta} + \frac{1-\cos\theta}{1+\cos\theta}$$

- Compares very well (also in magnitude) with data from the LEP detector L3 at $\sqrt{s} = 183 \text{ GeV}, 189 \text{ GeV}$ see hep-ex/0002036.

- Can put limits of $\Lambda_{\pm} \gtrsim 300 \text{ GeV}$ on new interactions producing

$$\frac{d\sigma}{d\Omega} = \left(\frac{d\sigma}{d\Omega} \right)_{\text{QED}} \left(1 + \frac{s^2}{\Lambda^4} \frac{1}{4} \sin^2\theta \right)$$

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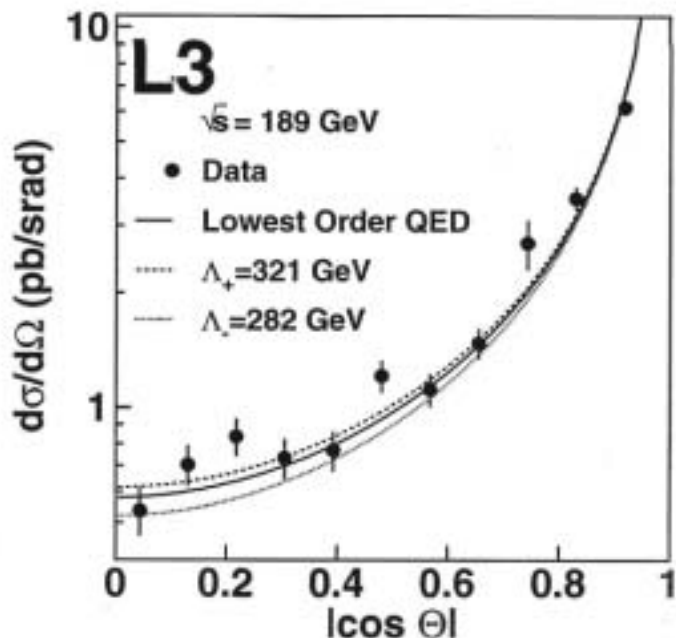
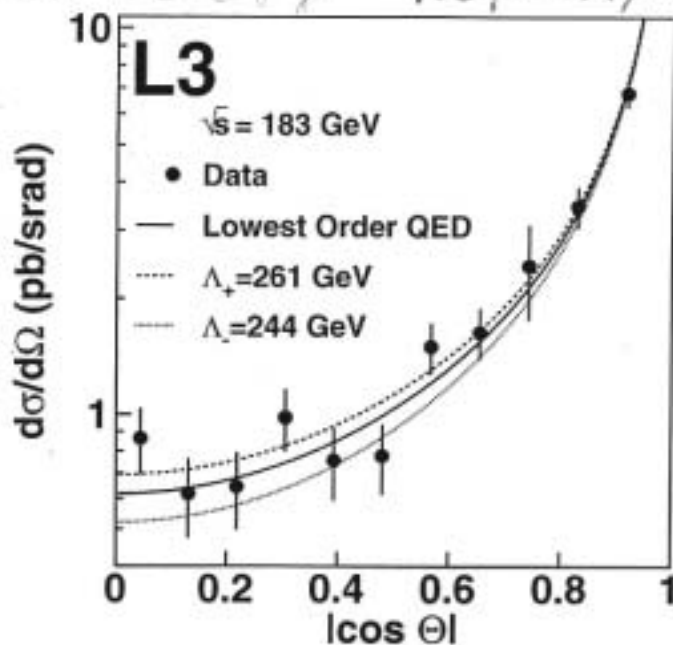
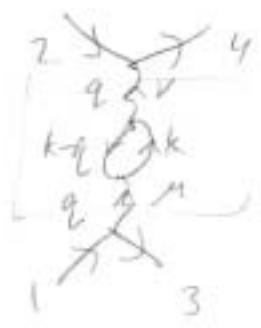


Figure 4: Differential cross section as a function of $\cos\theta$ for the process $e^+e^- \rightarrow \gamma\gamma(\gamma)$. The points show the measurements corrected for efficiency and additional photons. The solid line corresponds to the lowest order QED prediction. The dashed and dotted lines represent the limits obtained for deviations from QED, taking into account all the L3 data at centre-of-mass energies up to that presented in the corresponding plot.

QED divergences & renormalization

Consider 1-loop correction to $e\gamma$ scattering, in particular



"vacuum polarization" due to virtual e^+e^- pairs

Neglecting m_e ,

$$\mathcal{M} = +ie^2 \bar{u}_3 \gamma_\mu u_1 \left[\int \frac{d^4k}{(2\pi)^4} \frac{\text{Tr}[\gamma^\mu \not{k} \gamma^\nu (\not{k}-\not{q})]}{k^2 (k-q)^2} \right] \bar{u}_4 \gamma_\nu u_2$$

$\equiv I^{\mu\nu}(q)$

Note that $I^{\mu\nu}(q) = e^2 \int \frac{d^4k}{(2\pi)^4} \frac{\text{Tr}[\gamma^\mu \not{k} \gamma^\nu (\not{k}-\not{q})]}{k^2 (k-q)^2}$

Similarly, $q_\nu I^{\mu\nu}(q) = 0$

$$= 4i \int \frac{d^4k}{(2\pi)^4} \left[\frac{(k-q)^\nu}{(k-q)^2} - \frac{k^\nu}{k^2} \right] = 0$$

at least, it should be if regulator is OK

$\therefore I^{\mu\nu}(q) = (q^2 \delta^{\mu\nu} - q^\mu q^\nu) I(q^2)$

only this term contributes to \mathcal{M}

this term is easier to work out

Feynman's trick to combine denominators.

$$\frac{1}{k^2 (k-q)^2} = \frac{1}{q^2 \ell \cdot k} \left[\frac{1}{k^2} - \frac{1}{(k-q)^2} \right] = \int_0^1 dx \frac{1}{[k^2 - 2x k \cdot q + x q^2]^2}$$

$$= \int_0^1 dx \frac{1}{(\ell^2 + x(1-x)q^2)^2} \quad \text{where } \begin{cases} \ell = k - xq \\ k = \ell + xq \end{cases}$$

So,

$$I^{\mu\nu} = e^2 \int \frac{d^4 l}{(2\pi)^4} \int_0^1 dx \frac{\text{Tr} [\gamma^\mu (\not{q} + x \not{A}) \gamma^\nu (\not{q} - (1-x) \not{A})]}{[l^2 + x(1-x)q^2]^2}$$

$$\frac{1}{4} \text{Tr}[\dots] = (l^\mu + xq^\mu) \cdot (l^\nu - (1-x)q^\nu) + \cancel{\mu \leftrightarrow \nu} - g^{\mu\nu} (l^2 + (2x-1)l \cdot \cancel{\epsilon} - x(1-x)q^2)$$

$$= 2l^\mu l^\nu - g^{\mu\nu} l^2 - 2x(1-x)q^\mu q^\nu + g^{\mu\nu} x(1-x)q^2 + \text{[linear in } l^\mu, l^\nu]$$

integrate to 0 due to $l^\mu \leftrightarrow -l^\mu$ symmetry

$$\int \frac{d^4 l}{(2\pi)^4} \frac{l^\mu l^\nu}{[l^2 + x(1-x)q^2]^2} = \cancel{g^{\mu\nu}} \cdot f(l^2)$$

We just want the $q^\mu q^\nu$ terms

$$\Rightarrow I(q^2) = e^2 \cdot 2 \cdot 4 \int \frac{d^4 l}{(2\pi)^4} \int_0^1 dx \frac{x(1-x)}{[l^2 + x(1-x)q^2]^2}$$

"UV cutoff"

Now $\int \frac{d^4 l}{(2\pi)^4} \frac{1}{(l^2 + x(1-x)q^2)^2} \approx \frac{\int d^4 l}{(2\pi)^4} \int_0^M \frac{dl l^3}{l^4}$ (vol. of 25-d sphere)

$$\approx \frac{2\pi^2}{(2\pi)^4} \cdot \frac{1}{2} \int_0^{M^2} \frac{dl^2}{l^2} = \frac{1}{16\pi^2} (\ln M^2 + \text{finite})$$

$$\text{So, } I(q^2) = \frac{e^2}{2\pi^2} \ln M^2 \underbrace{\int_0^1 dx x(1-x)}_{\frac{1}{2} - \frac{1}{3} = \frac{1}{6}} + \text{finite}$$

$$I(q^2) = \frac{e^2}{12\pi^2} \ln M^2 + \text{finite}$$

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calling our original $e \rightarrow e_0(M)$
So, finally, ~~the bare charge~~ "the bare charge"

$$\mu = -e_0^2 \bar{u}_3 \gamma^\mu u_1 \frac{g_{\mu\nu}}{q^2} \left\{ 1 - \frac{e_0^2}{12\pi^2} \left[\ln\left(\frac{M^2}{m_e^2}\right) = f\left(\frac{q^2}{m_e^2}\right) \right] \right\} \bar{u}_4 \gamma^\nu u_2$$

To remove the divergence, we define the renormalized charge

$$e_R = e_0^{(M)} \sqrt{1 - \frac{e_0^2}{12\pi^2} \ln\left(\frac{M^2}{m_e^2}\right)}$$

such that e_R is finite as $M \rightarrow \infty$

Then

$$\mu_{(ren)} = -e_R^2 \bar{u}_3 \gamma^\mu u_1 \frac{g_{\mu\nu}}{q^2} \left\{ 1 + \frac{e_R^2}{12\pi^2} f\left(-\frac{q^2}{m_e^2}\right) \right\} \bar{u}_4 \gamma^\nu u_2$$

$\mu_{(ren)}$ is finite as $M \rightarrow \infty$

For $q^2 \gg m_e^2$, $f\left(-\frac{q^2}{m_e^2}\right) \approx \ln\left(\frac{q^2}{m_e^2}\right)$

(to cancel m_e dependence in $\left[\ln\left(\frac{M^2}{m_e^2}\right) - f\left(-\frac{q^2}{m_e^2}\right) \right]$)

We can ~~cancel~~ ^{absorb} this dependence ~~from~~ ⁱⁿ μ too, in terms of a running coupling:

$$e_R(q^2) = e_R(0) \sqrt{1 + \frac{e_R(0)^2}{12\pi^2} f\left(-\frac{q^2}{m_e^2}\right)}$$

$$\text{or } \alpha(q^2) = \alpha(0) \left[1 + \frac{\alpha(0)}{3\pi} f\left(-\frac{q^2}{m_e^2}\right) \right] \approx \alpha(0) \left[1 + \frac{\alpha(0)}{3\pi} \ln\left(\frac{q^2}{m_e^2}\right) \right] \quad |q| \gg m_e$$

α increases ^{logarithmically} as q increases

