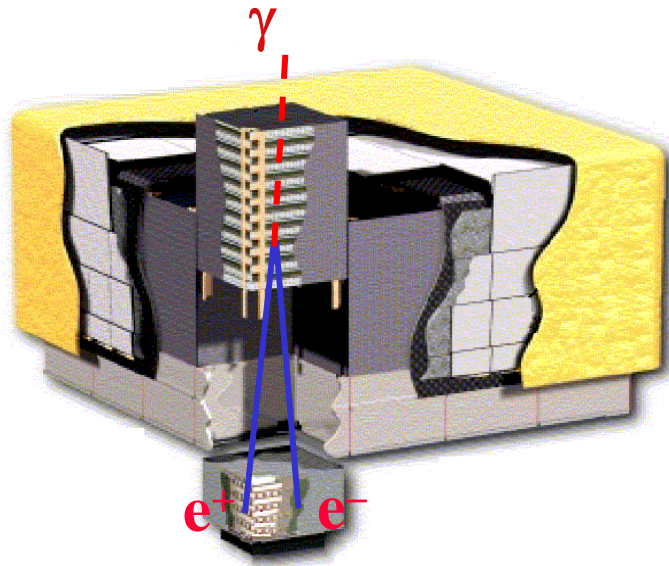




Project Overview and Collaboration Status



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GLAST LAT Principal Investigator and Spokesperson
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Recent Events

- ▶ **flight hardware now arriving at SLAC for instrument integration**
 - *2 tracker modules, 7 calorimeter modules, 2 tower electronics modules, instrument Grid*

- ▶ **GLAST rebaseline: necessary to address schedule and cost growth**
 - *Feb 24 discussion with NASA Headquarters about possible LAT descope; no actions directed*

- ▶ **Collaboration meeting, scheduled for March 8-10, cancelled due to priority placed on flight hardware manufacture and delivery schedule – next meeting scheduled for early September**

- ▶ **I&T Data Analysis Workshop held at SLAC, March 10, 2005**

- ▶ **Instrument Science Operations Center planning well underway – ISOC Manager, Rob Cameron**

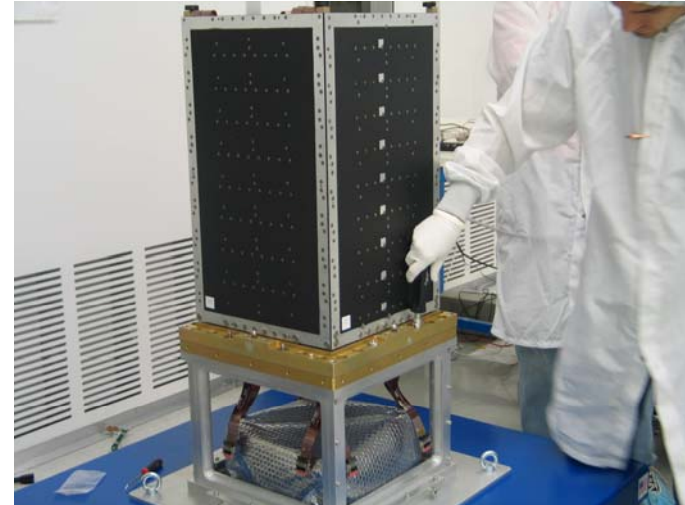
- ▶ **Science Analysis Planning underway**
 - *science analysis plan formulated and science group coordinators identified*
 - *preparations for DC-2*
 - *proposal for 1st year data release to community prepared*



Flight Hardware Integration



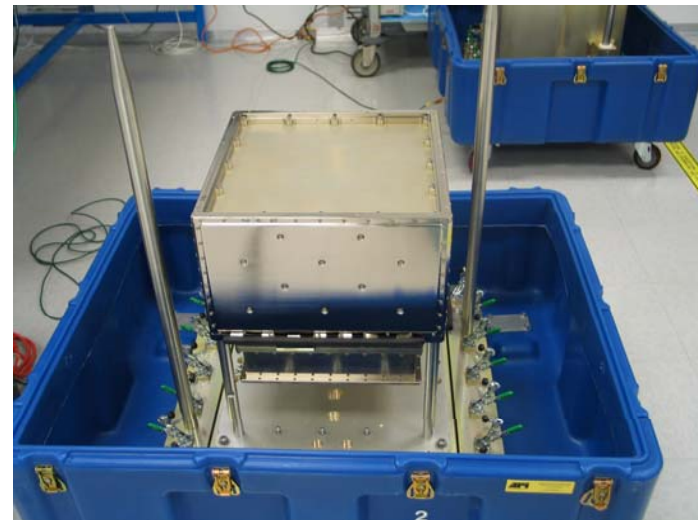
Preparation of flight grid for TCS integration



Flight Tracker in Cleanroom at SLAC



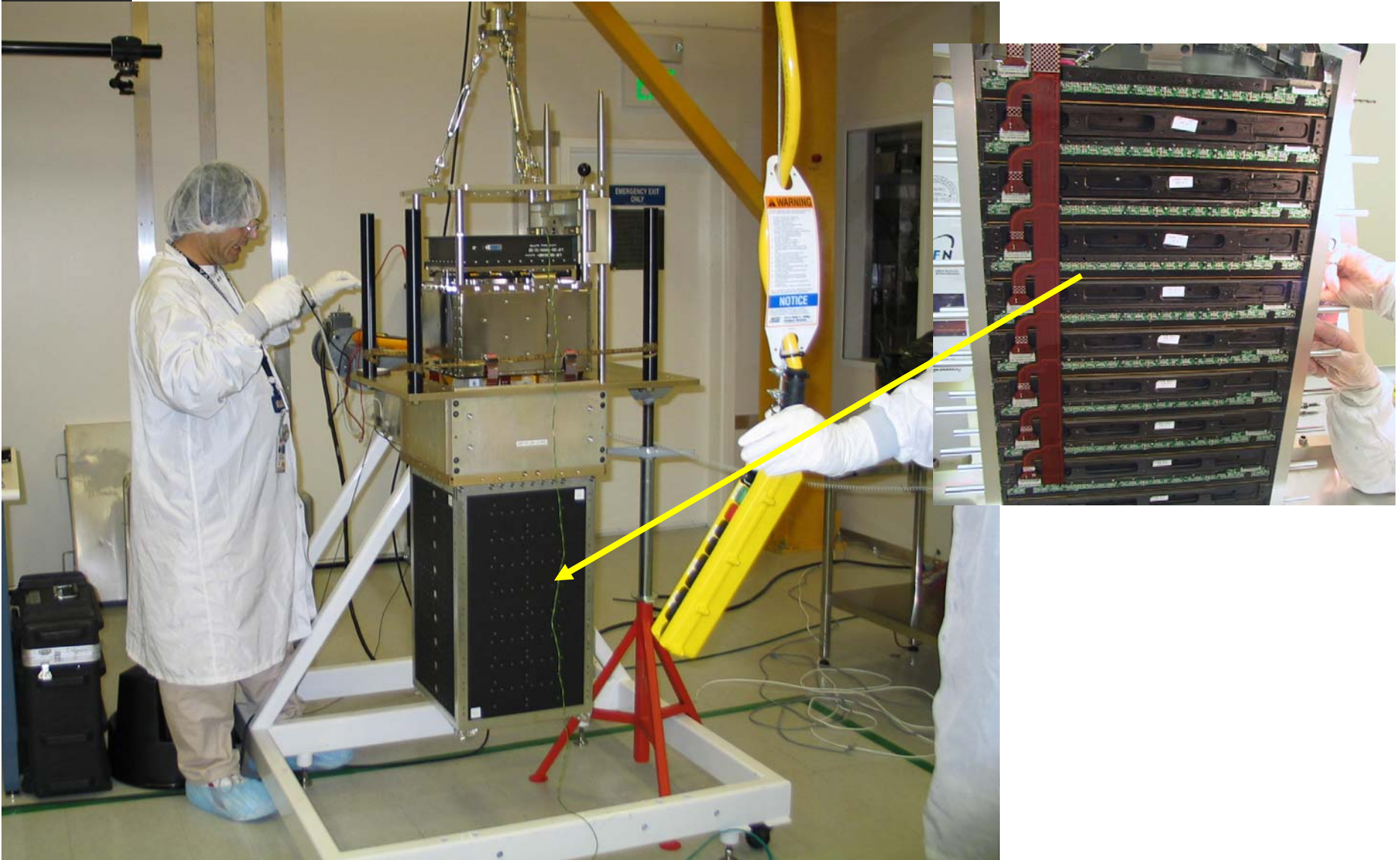
LAT Integration stand with PAP ready for proof test



Flight Calorimeter



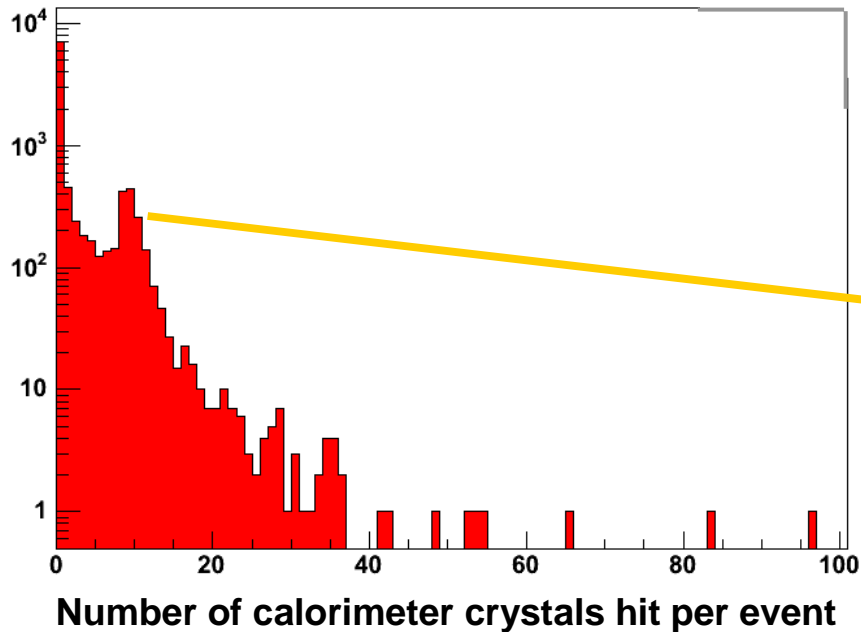
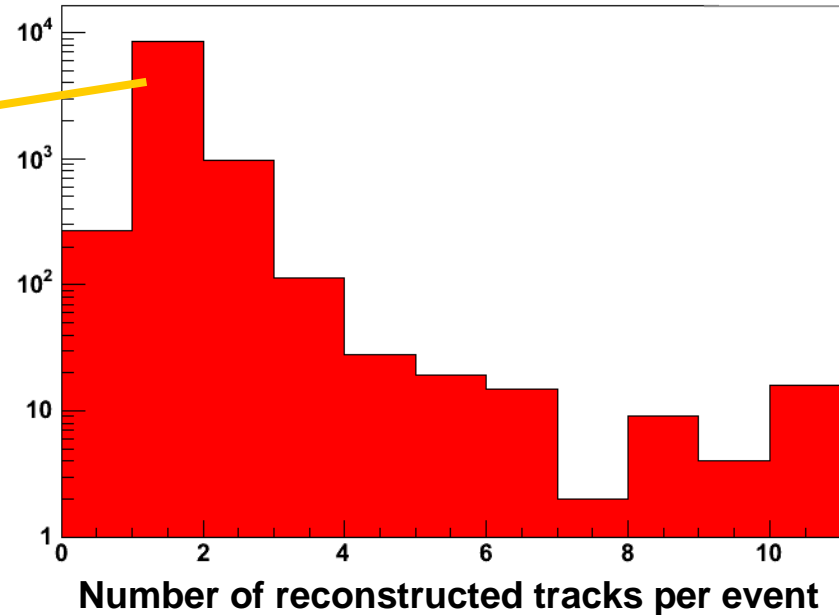
First Flight Tower in I&T





First Integrated Tower – Cosmic Ray Data

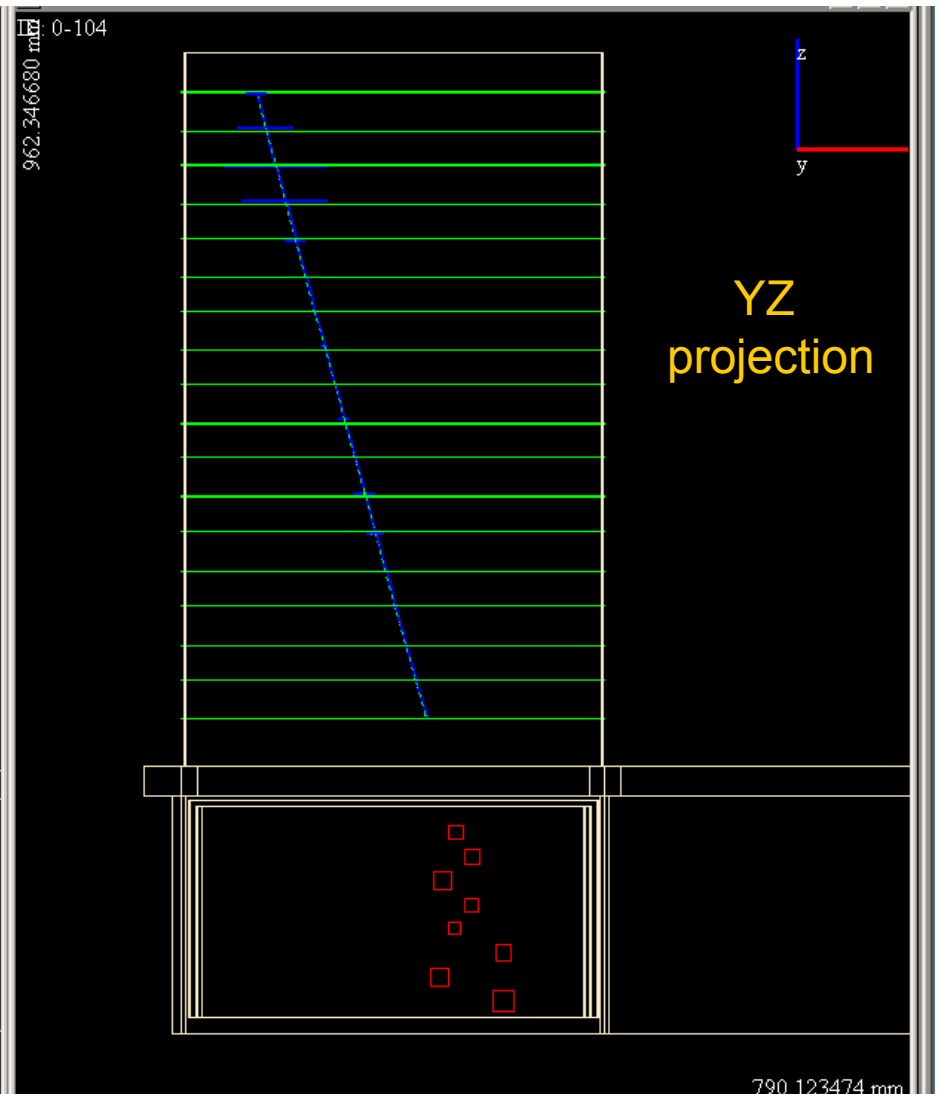
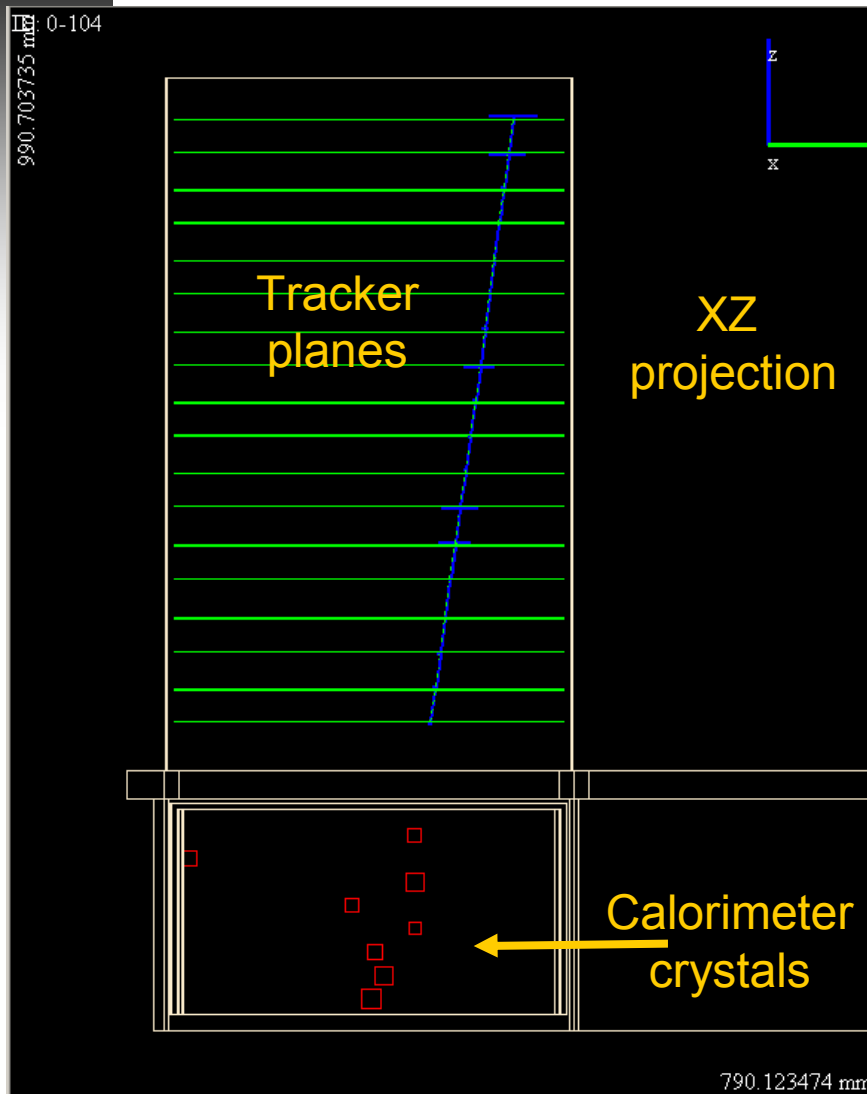
Most of the cosmic ray (muon) events are expected to have single tracks



Most of the cosmic ray (muon) events with on-axis incidence are expected to hit about 8 CAL crystals

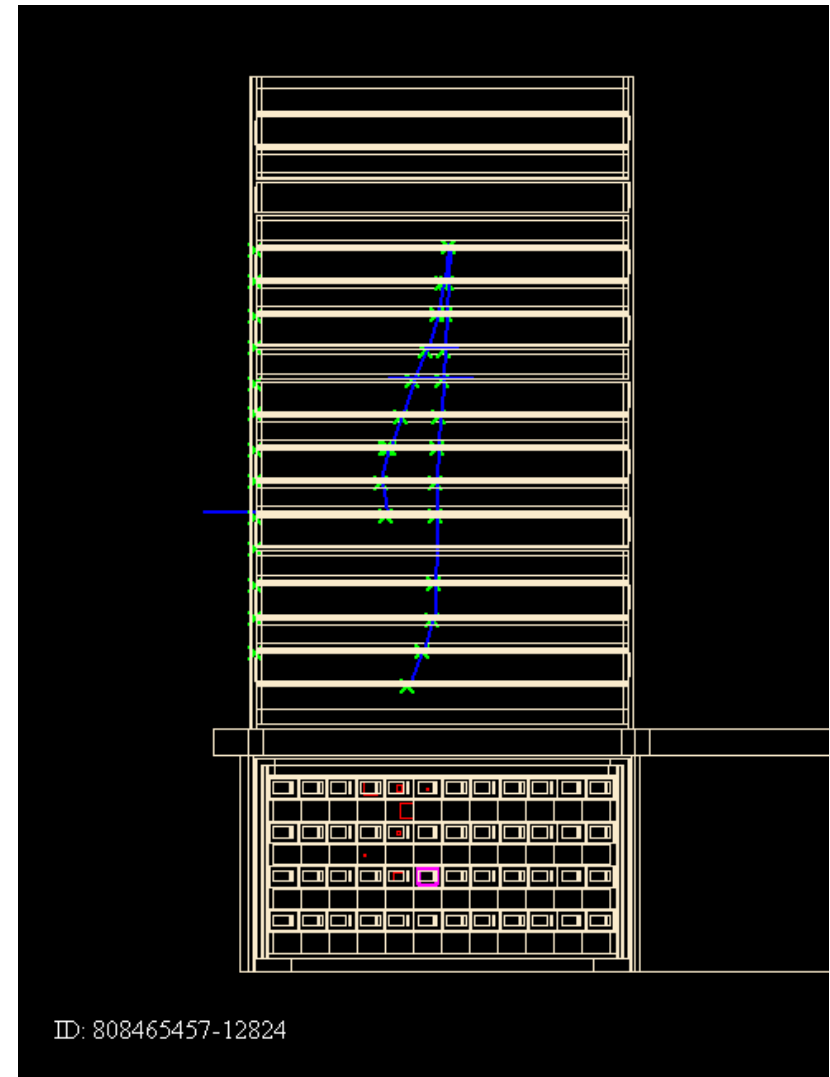
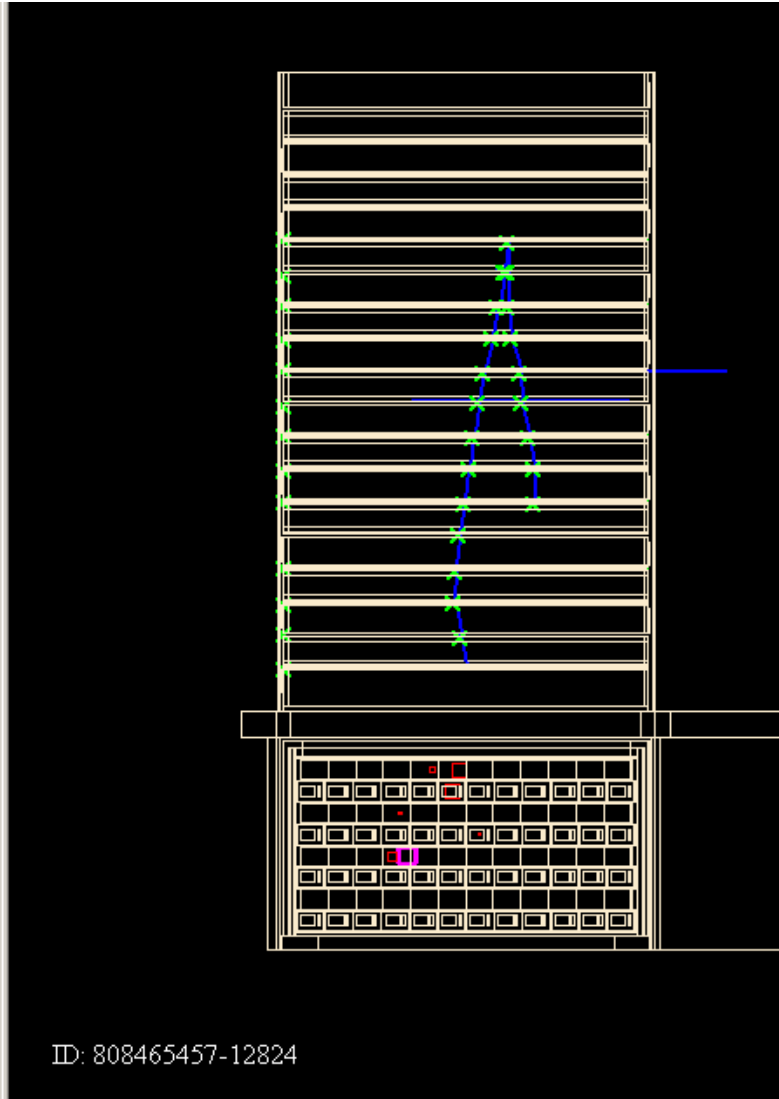


First Integrated Tower – Muon Candidate Event





Gamma-ray Candidate Event





GLAST Discovery Reach

► ***Many opportunities for exciting discoveries:***

- *determine the origin(s) of the high-energy extragalactic diffuse background*
- *measure extragalactic background light to $z > 3$*
- *detect γ -ray emission from clusters of galaxies; cosmic-ray acceleration on large scales*
- *detect γ -rays from Ultra-Luminous Infrared Galaxies; cosmic ray acceleration efficiency and star formation rate*
- *detect high-latitude Galactic Inverse-Compton emission and thereby measure TeV-scale CR electrons in the Galaxy*
- *the unknown!*

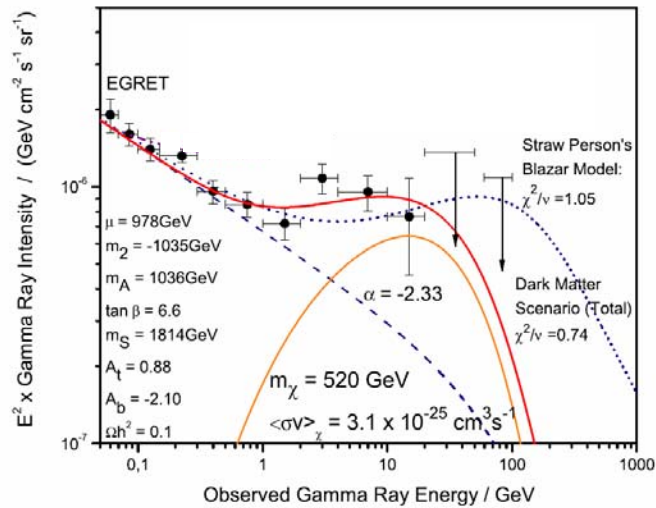


Cosmology: Origin of Extragalactic Diffuse Radiation

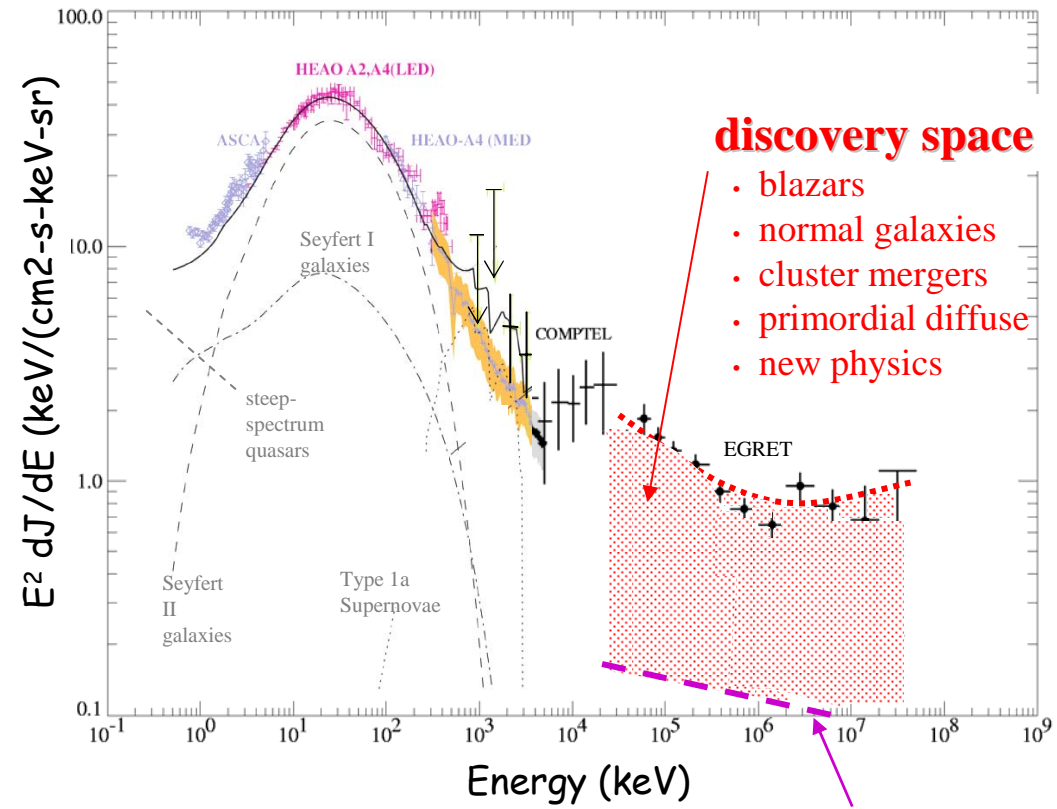
- ▶ **origin is a mystery; either sources there for GLAST to resolve (and study!) OR there is a truly diffuse flux from the early Universe**

EGRET constrains blazars to be > 25% of diffuse;

annihilation of cosmological neutralinos has, in principle, a distinctive spectral signature



from Elasser & Mannheim, astro-ph/0405235



discovery space

- blazars
- normal galaxies
- cluster mergers
- primordial diffuse
- new physics

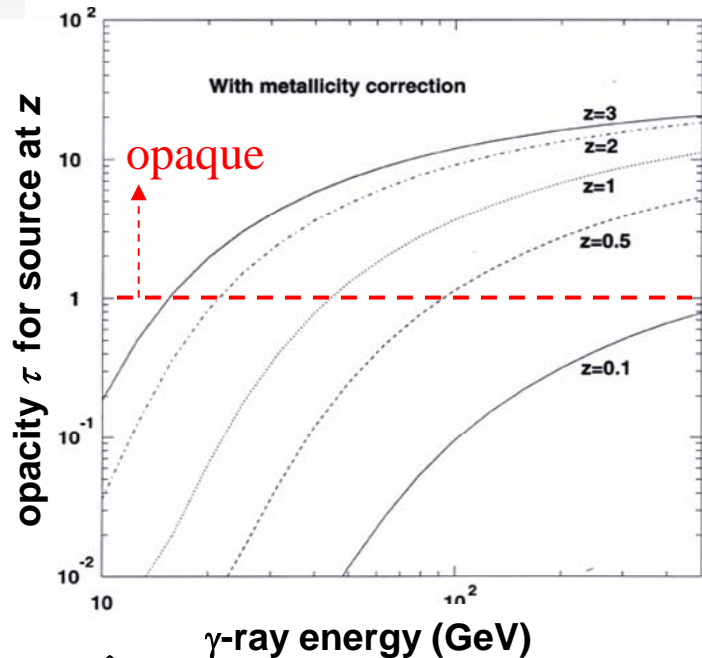
LAT baseline background limit

Unique science for GLAST

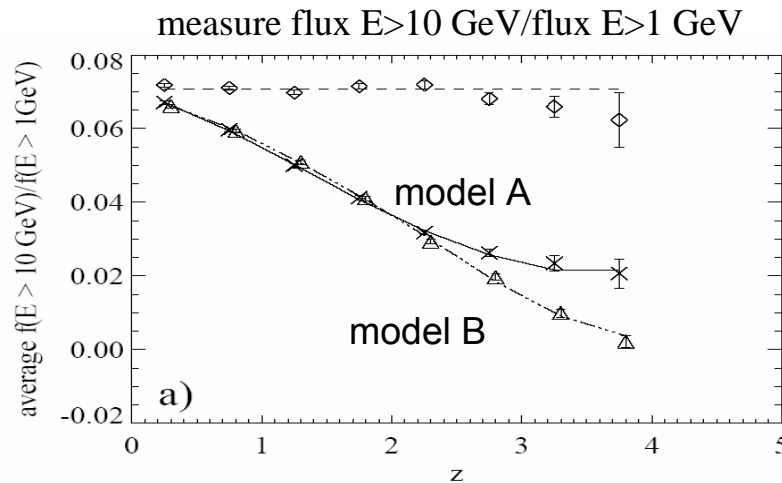


Cosmology: Probing Extragalactic Background Light

- ▶ **High-redshift ($z > 3$) blazar detections important for visible-UV EBL studies and blazar/galaxy evolution**
 - **Number of blazar detections with spectral measurements above 10 GeV important**



↑ No significant attenuation below 10 GeV



number of sources in bins of z
luminosity-function dependent (important science for GLAST!)

	# blazars with spectra $z > 3$
$E > 1 \text{ GeV}$	62
$E > 10 \text{ GeV}$	12

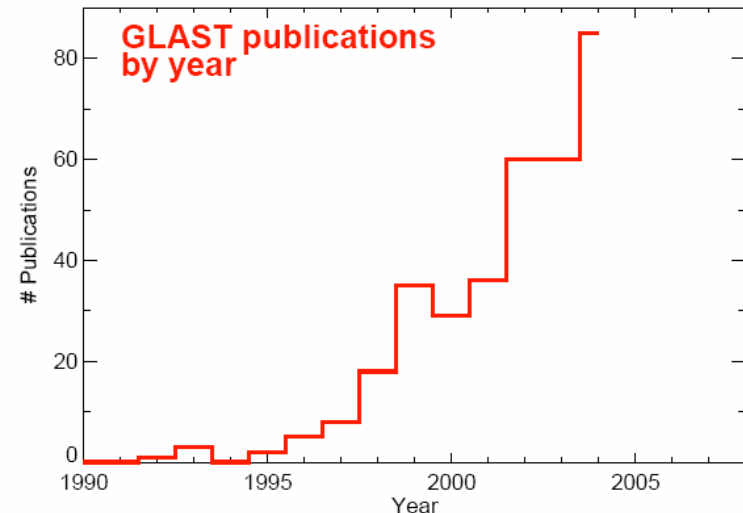


Broad Science Community Involvement

- ▶ **At all mission levels, including Conferences, Workshops.**
- ▶ **Multiwavelength survey initiatives underway**
 - VLA monitoring for a large sample of flat-spectrum, compact sources (Ulvestad, et al.)
 - VLBA Imaging and Polarimetry Survey proposed to obtain a set of reference images for 1000 potential LAT sources in advance of GLAST launch (Taylor, et al.).
 - Northern and Southern hemisphere radio pulsar timing campaigns in support of GLAST mission (Thorsett, et al. ; Manchester, et al.)
 - optical monitoring of LAT AGN
- ▶ **Broad and growing interest**

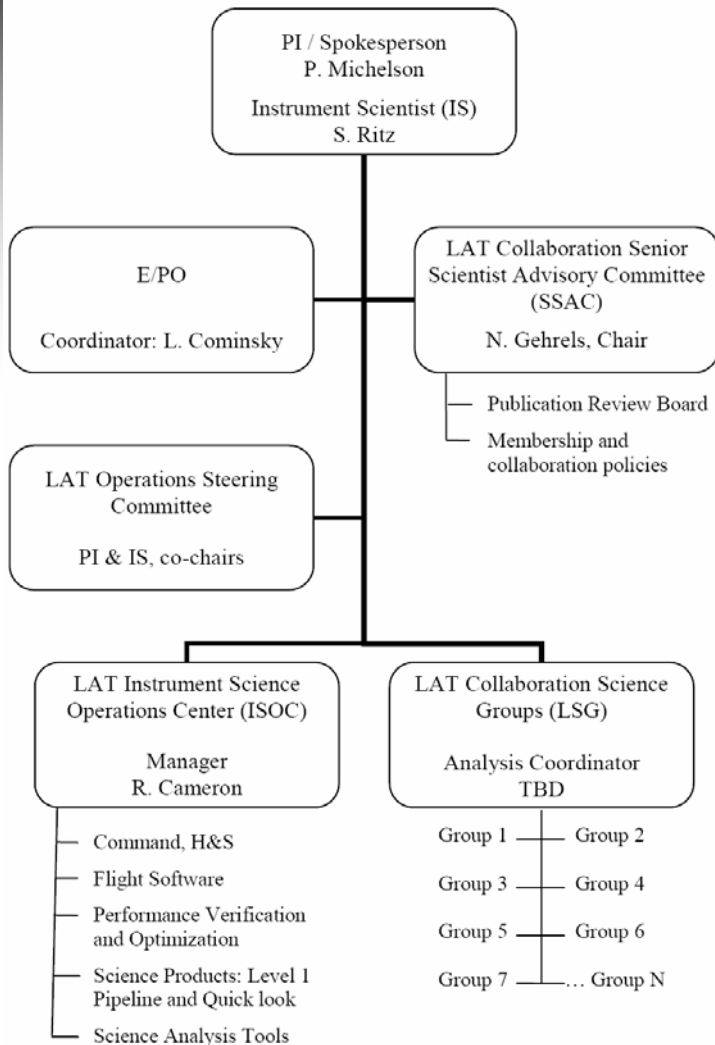
Publications & Conference Proceedings

		pre-launch	total
COS-B	1975-1982	11	481
EGRET	1991-2000	53	1212
GLAST	2007-	342	???





Operations Phase LAT Organization Chart



► Instrument Science Ops Center (ISOC)

- responsible for instrument operation (monitoring, configuration management, flight software, calibration, science analysis software, etc) and operation of Level 1 Data Pipeline
- performance analysis and tools development are coordinated across the Collaboration by the ISOC

► LAT Collaboration Science Groups

- responsible for collaboration's analysis and extraction of science results from LAT data; including galactic diffuse model and source catalogs
- Science Groups plus Multi- λ Coordination Working Group
- each science group to have 2 coordinators; with one resident at Stanford (in particular during 1st year)
- day-to-day efforts coordinated by Analysis Coordinator, also resident at Stanford-SLAC



GLAST LAT Science Groups

Group	Principal Responsibilities
1. Catalogs	Prepare comprehensive LAT catalog for publication, after 1, 2, and 5 years (nominal), including cross-identifications when known.
2. Galactic Diffuse and Molecular Clouds	Generate Galactic diffuse model for delivery to community (GSSC) and provide periodic updates; study mass tracers and CR density in individual clouds
3. Extragalactic Diffuse	Analysis of the residual isotropic gamma-ray emission (after accounting for LAT sources and Galactic diffuse) including its spectra and angular fluctuations; assess contributions from various source classes
4. Blazars & Other AGNs	Analysis of spectra and time variability of LAT-detected blazars and other types of AGNs; modeling of spectra (e.g. SSC model) including effects of EBL; study evolution with redshift
5. Other Galaxies	Search for emission from other galaxies, particularly nearby galaxies such as M31, ULRGs, and from clusters of galaxies; determine source spectra and/or set upper limits
6. Pulsars, SNRs & Plerions	Construct phase-resolved "light curves" and spectra for LAT-detected radio pulsars; search for periodic emission in other sources, including unidentified sources; identify SNRs and plerions and model spectra of these sources
7. Unidentified Sources & Population Studies	Coordinate deep multiwavelength studies of the error boxes of the unidentified GLAST sources and correlate with potential counterpart populations; determine the heterogeneity of the population and correlate with source characteristics (e.g. time variability; spectral index, etc.)
8. Dark Matter and Exotic Physics	Search for signatures of dark matter, extra dimensions (e.g. KK graviton decay), quantum gravity dispersion effects, etc.
9. Gamma-Ray Bursts	Construct LAT GRB catalogs; determine burst spectra, including time evolution; coordinate joint analysis of bursts with GBM
10. Sources in the Solar System	Solar flares; lunar emission; ...
11. Calibration and Analysis Methods	Pre-flight performance of the LAT and on-orbit calibration and performance (1) Monte Carlo model of LAT; event reconstruction and background rejection methods; (2) astrophysical data analysis methods (e.g. likelihood analysis; time series analysis; source spectral analysis, cross-correlation methods, etc.)



Proposed LAT Year-1 Data Release Plan

► **consistent with NASA AO 99-OSS-03:**

- *“At all times, including during the first twelve months of science operations, the data from transient sources discovered or detected by GLAST will immediately be made publicly available.”*
 - *operational definition of transient source: any source for which a significant change in source flux is detected on a timescale sufficiently short that rapid follow-up multi-wavelength observations are warranted (e.g. GRBs, a significant flare from a blazar, a solar flare, etc.)*

- *“During the first twelve months of science operations, data from specific sources of interest to qualified individual researchers will be made available...”*
 - *Data products on specific sources of interest will be released, by the LAT team, periodically via a publicly assessable web site.*



Proposed LAT Year-1 Data Release Plan

- ▶ ***during Phase 1 (sky survey & verification phase), event reconstruction & background rejection cuts will be refined in several stages***
 - *each modification of algorithms will likely require reprocessing of level-0 to level-1 data;*
 - *expect frequency of reprocessing to be relatively high and effects on level-1 data base to be more significant in Phase 1 than during subsequent phases;*

- ▶ ***during Phase 1: release high-level data on transients and monitored sources consisting of flux (fluence) on various timescales, spectra, source position, and errors (including estimate of systematic errors) for all of these quantities;***
 - *no release in phase 1 of individual reconstructed photon events.*



Some specifics

▶ **GRBs:**

- *on-board detections released “immediately”*
- *refinement of parameters (e.g. source position, fluence in several energy bands)with subsequent processing on the ground; release information as soon as available.*
- *for GRBs detected by GBM but not by LAT, a LAT upper limit will be released.*

▶ **Sources of Interest:**

- *monitor and regularly release data, via the web, on list of sources of interest*
- *update data weekly and include fluxes (daily to weekly averages), estimates of the spectra (weekly averages), position, and errors*
- *expected latency between receipt of level-0 data at ISOC and availability of processed high-level data will improve with time. Nominal schedule is*
 - *first month of Phase 1; 2 weeks*
 - *month 2-6; within 3 days*
 - *months 6-12; within 1 day*
 - *months 13-N; within 1 day (goal of 6 hours)*



Conclusion

- ▶ ***GLAST flight hardware in production; BEAUTIFUL data from first integrated flight tower***
- ▶ ***Collaboration is actively engaged in the production, integration, and testing of flight hardware;***
- ▶ ***also continuing effort to develop science analysis tools; DC-2 preparations underway***
- ▶ ***working with broader community to support / initiate multiwavelength observations and surveys needed for GLAST science***
- ▶ ***science planning will be major focus of September Collaboration meeting***