

7. Applications of Superspace

Now that we have developed the superspace formalism, which gives actions and other expressions that are manifestly $N=1$ supersymmetric, I would like to describe some applications.

① Superspace perturbation theory.

It is possible to develop Feynman perturbation theory directly from the superspace action. I will not describe in detail how to do this (there is an extensive discussion in Wess + Bagger) but I would like to quote some results. The propagators in this formalism are

$$\frac{\Phi}{\leftarrow} = \langle \Phi \Phi^* \rangle = i \frac{\frac{1}{16} \bar{D}^2 D^2}{p^2} = i \times \begin{pmatrix} \text{projector onto } \theta \\ \text{chiral superfield} \end{pmatrix}$$

$$\overset{V}{\leftarrow} = \langle V V \rangle = -i \frac{1}{p^2} \quad (\text{in a Feynman sense})$$


The vertices are localized in superspace. Note that, while the chiral field propagator has components that are fairly conventional, the vector field propagator leads to compound Feynman rules such as

$$\langle \phi D \rangle = -\frac{i}{p^2} \quad \langle \psi_\alpha \partial_\beta \rangle = -\frac{i \epsilon_{\alpha\beta}}{p^2}$$

so the whole field content of the vector superfield becomes involved

in the perturbation theory.

On the other hand, the results of this perturbation theory must be manifestly supersymmetric and so must exhibit the SUSY non-renormalization theorems. Here are some examples:

a) All vacuum diagrams (e.g. ) are proportional to $\int d^4\theta [\theta\text{-indep}] = 0$. Thus, all perturbative contributions to the vacuum energy vanish diagram by diagram.

b) All contributions to the effective Lagrangian for chiral superfields are of the form

$$\int d^4x \int d^4\theta (\text{func of } \Phi, \Phi^*)$$

Notice that all such interactions, when reduced to components, have derivatives. The structure

$$\int d^2\theta (\text{func of } \Phi)$$

is not generated by supergraphs. So, there are no perturbative corrections to the superpotential.

c.) From the argument, you might think that it is also not possible to generate the structure

$$\int d^2\theta W^2$$

If we rescale $gV \rightarrow V$ so that g is removed from all couplings, the kinetic term of the gauge field is

$$\int d^3\theta \frac{1}{4g^2} W^2 + h.c.$$

So the radiative corrections of the form $\int d^3\theta \cdot \Delta \cdot W^2$ generate coupling constant renormalization. In a fixed gauge, this structure can be rewritten as

$$\int d^4\theta \Delta \cdot V(p)^2 V(-p)$$

and can be generated by supergraphs. However, in a well-chosen regulator, this contribution appears only at 1-loop order.

Later in this lecture, I will make this more clear by connecting this result to the axial vector anomaly.

d.) In 1-loop order, corrections to the Fayet-Iliopoulos term

$$\int d^4x \kappa D = \int d^4x d^4\theta \kappa V$$

are proportional to the sum of $U(1)$ charges of the matter fields:

$$\text{loop} \propto \text{tr } Q$$

So if $\text{tr } Q = 0$, this correction vanishes to 1-loop order. It is possible to show that, in that case, the correction to κ vanishes to all orders.

② Soft SUSY-breaking perturbations

In a th_y with exact SUSY, there are no quadratically divergent corrections to parameters in the Lagrangian. Now, the real world is not supersymmetric, so it is worth asking whether we can add perturbations to \mathcal{L} while preserving this property. From ordinary gauge field th_y, we might expect that perturbations to \mathcal{L} do not disturb the symmetry structure if the terms we add have dimension < 4 . However, it is not so hard to find an operator with dimension < 4 that leads to quadratic divergences. Consider

$$\delta\mathcal{L} = \int d^4x \, g \phi^\dagger \phi^2 + h.c.$$

Then $\int \frac{d^4k}{k^2} \rightarrow$  leads to $\int d^4x \, (g \Lambda^2) \phi$

that is, we induce a correction to the potential that is quadratically divergent and can shift $\langle \phi \rangle$ by a divergent amount. We would like to know if this is a generic situation, or if there are specific operators we can add that avoid this difficulty.

Grisaru and Girardello analyzed this problem in superspace. Imagine that we write a supersymmetric

Lagrangian that contains some extra chiral field Φ_0 or vector field V_0 . Perturbate this with this Lagrangian will never generate quadratically divergent corrections to \mathcal{L} .

Now replace

$$\Phi_0 \rightarrow \theta^2 F \quad \text{or} \quad V_0 \rightarrow \theta \bar{\theta}^2 d$$

These replacements break SUSY. But the terms that result still cannot generate quadratic divergences.

Let's apply this to various cases.

$$\int d^3\theta (\Phi_0 \Phi, \Phi_0 \Phi^2, \Phi_0 \Phi^3) \rightarrow (f \cdot \Phi, f \Phi^2, f \cdot \Phi^3)$$

$$\int d^4\theta \Phi^\dagger \Phi V_0 \rightarrow d \Phi^\dagger \Phi$$

so the perturbations of these forms are "soft" in the SUSY sense that they do not generate quadratic divergences.

Also, the Lagrangian

$$\int d^3\theta \Phi_0 W^2$$

is supergauge specifically if Φ_0 is a chiral superfield;

this generates

$$f \lambda^T a \lambda^a$$

a supersymmetric mass term. So, a SL of the form

$$\begin{aligned} \mathcal{L} = & [(\text{holomorphic polynomial in } \phi_k) + \text{h.c.}] \\ & + \\ & (\text{scalar mass term} \quad m_{kl} \phi_k^* \phi_l) \\ & + \\ & (\text{gaugino mass term} \quad m_a \lambda^a \epsilon \lambda^a) \end{aligned}$$

is soft in the sense that we require. Of possible SUSY-breaking perturbations with dimension ≤ 3 , the only terms that do not appear on the list above are

$$\int \phi^* \phi^2 + \text{h.c.} \quad m \psi^T \epsilon \psi$$

We have already seen that the first of these interactions generates quadratic divergences. For the matter fermion mass, if we link it to a superpotential vertex:

$$\text{Diagram: a circle with a star on top and a vertical line with an arrow pointing down from the bottom.} \quad \sim m \lambda \left(\int \frac{d^4 k}{k^2} \right) \phi \sim m \lambda \Lambda^2 \phi$$

we can also generate a divergence. So the list at the top of this page is complete.

③ $N=1$ SUSY nonlinear sigma models

Up to now, we have only considered the simplest kinetic term for the chiral superfield

$$\int d^4\theta \Phi^* \Phi$$

From superspace, it is clear that the most general 2-derivative Lagrangian for a set of chiral superfields is

$$\int d^4\theta \mathcal{K}(\phi_k, \phi_k^*)$$

Let's now investigate the physics of this more general interaction.

We need to compute $[\mathcal{K}(\phi, \phi^*)]_{\theta^2 \bar{\theta}^2}$, so start expanding:

$$\begin{aligned} \mathcal{K} &= \mathcal{K}(\phi, \phi^*) + \frac{\partial \mathcal{K}}{\partial \phi_k} (\sqrt{2} \theta^T c \psi_k + \dots) \\ &\quad + \frac{1}{2} \frac{\partial^2 \mathcal{K}}{\partial \phi_k \partial \phi_l} (\sqrt{2} \theta^T c \psi_k + \dots) (\sqrt{2} \theta^T c \psi_l + \dots) \\ &\quad + \dots \end{aligned}$$

Abbreviate:

$$\frac{\partial \mathcal{K}}{\partial \phi_k} = \mathcal{K}_{,k} \quad \frac{\partial \mathcal{K}}{\partial \phi_k^*} = \mathcal{K}_{,k^*} \quad \text{etc.}$$

the terms involving two derivatives on ϕ or ϕ^* are

$$\begin{aligned}
& K_{,k} \left(-\frac{1}{4} \theta^2 \bar{\theta}^2 \partial^2 \phi_k \right) + K_{,k} \left(-\frac{1}{4} \theta^2 \bar{\theta}^2 \partial^2 \phi_k^* \right) \\
& + \frac{1}{2} K_{,kl} \left(i \bar{\theta} \bar{\sigma}^a \theta \partial_a \phi_k \right) \left(i \bar{\theta} \bar{\sigma}^b \theta \partial_b \phi_l \right) + h.c. \\
& + K_{,k\bar{l}} \left(-i \bar{\theta} \sigma^a \theta \partial_a \phi_l^* \right) \left(i \bar{\theta} \bar{\sigma}^b \theta \partial_b \phi_k \right) \\
= & \theta^2 \bar{\theta}^2 \left\{ -\frac{1}{4} K_{,k} \partial^2 \phi_k - \frac{1}{4} K_{,k} \theta^2 \phi_k^* \right. \\
& \quad \left. + \frac{1}{4} K_{,kl} \partial_a \phi_k \partial^a \phi_l - \frac{1}{4} K_{,k\bar{l}} \partial_a \phi_k^* \partial^a \phi_l^* \right. \\
& \quad \left. + \frac{1}{2} K_{,k\bar{l}} \partial_a \phi_l^* \partial^a \phi_k \right\}
\end{aligned}$$

Note that

$$\begin{aligned}
\int d^4x \left(-\frac{1}{4} K_{,k} \partial^2 \phi_k \right) &= \int d^4x \frac{1}{4} \partial_a (K_{,k}) \partial^a \phi_k \\
&= \int d^4x \left(\frac{1}{4} K_{,kl} \partial_a \phi_l \partial^a \phi_k + \frac{1}{4} K_{,k\bar{l}} \partial_a \phi_l^* \partial^a \phi_k \right)
\end{aligned}$$

in which the first term cancels the 2nd line above and the second term adds to the third line to give finally:

$$\theta^2 \bar{\theta}^2 K_{,k\bar{l}} \partial_a \phi_l^* \partial^a \phi_k$$

It is not surprising that the terms with $\partial_k \phi \partial_l \phi$ cancel. If K were a function of Φ but not Φ^* , we would have

$$\int d^4\theta F(\Phi) = \int d^2\theta \left(-\frac{1}{4} \bar{D}^2 \right) F(\Phi) = 0$$

so any nonzero contributions to the kinetic term must involve both ϕ^* and ϕ .

Following our analysis of $\int d^4\theta \Phi^* \Phi$, we can find very quickly the form

$$\int d^4\theta \mathcal{K}(\Phi, \Phi^*) = K, \bar{K}_\ell (\partial_a \phi_k^+ \partial^a \phi_\ell + \frac{1}{2} \psi_k^+ i \bar{\sigma} \cdot \partial \psi_\ell - \frac{1}{2} \partial_a \psi_k^+ \bar{\sigma}^a \psi_\ell + F_k^* F_\ell) + (\text{terms at least cubic in } \partial\phi, \psi, F.)$$

Let's now gather these nonlinear terms:

$$\begin{aligned} & K, \bar{K}_\ell m (-\sqrt{2} \psi_k^+ c \bar{\theta}) (i \bar{\theta} \bar{\sigma}^a \theta \partial_a \phi_\ell) (\sqrt{2} \theta^T c \psi_m) \\ &= K, \bar{K}_\ell m (-2i) (-\frac{1}{2} \bar{\theta}^2) (\frac{1}{2} \theta^2) \psi_k^+ c c \bar{\theta}^a c c \partial_a \phi_\ell \psi_m \\ &= K, \bar{K}_\ell m \frac{i}{2} \theta^2 \bar{\theta}^2 \psi_k^+ (\bar{\sigma}^a \partial_a \phi_\ell) \psi_m \\ & K, \bar{K}_\ell m (\bar{\theta}^2 F_k^*) \frac{1}{2} (\sqrt{2} \theta^T c \psi_\ell) (\sqrt{2} \theta^T c \psi_m) \\ &= K, \bar{K}_\ell m \bar{\theta}^2 F_k^* \theta^2 \frac{1}{2} \psi_\ell^T c c c \psi_m \\ &= K, \bar{K}_\ell m (-\frac{1}{2}) \theta^2 \bar{\theta}^2 F_k^* \psi_\ell^T c \psi_m \end{aligned}$$

the conjugates of these terms also appear, as well as:

$$\begin{aligned} & K, \bar{K}_\ell m n \frac{1}{2} (-\sqrt{2} \psi_k^+ c \bar{\theta}) (-\sqrt{2} \bar{\theta}^T c \psi_\ell^*) \frac{1}{2} (\sqrt{2} \psi_m^T c \theta) (\sqrt{2} \theta c \psi_n) \\ &= K, \bar{K}_\ell m n (-\frac{1}{2} \bar{\theta}^2) (\frac{1}{2} \theta^2) \psi_k^+ c c c \psi_\ell^* \psi_m^T c c c \psi_n \\ &= K, \bar{K}_\ell m n (-\frac{1}{4}) \psi_k^+ \psi_\ell^* \psi_m^T c \psi_n \end{aligned}$$

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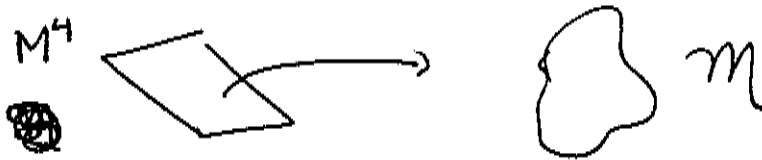
$$\begin{aligned}
 \int d^4\theta K(\Phi, \bar{\Phi}) = & K_{, \bar{k} \ell} \left\{ \partial_a \phi_k^* \partial^a \phi_\ell + \frac{i}{2} (\psi_k^+ \bar{\sigma}^a \psi_\ell - \partial_c \psi_k^+ \bar{\sigma}^a \psi_\ell) \right. \\
 & \left. + F_k^* F_\ell \right\} \\
 & + \frac{i}{2} K_{, \bar{k} \ell m} \psi_k^+ \bar{\sigma}^p \partial_p \phi_\ell \psi_m - \frac{i}{2} K_{, \bar{\ell} m k} \psi_k^+ \partial_a \phi_m^* \bar{\sigma}^a \psi_k \\
 & - \frac{1}{2} K_{\bar{k} \ell m} F_k^* \psi_\ell^T \psi_m + \frac{1}{2} K_{, \bar{k} \ell m} \psi_k^+ \psi_\ell^* F_m \\
 & - \frac{1}{4} K_{\bar{k} \ell m n} \psi_k^+ \psi_\ell^* \psi_m^T \psi_n
 \end{aligned}$$

Now, what is the significance of this structure?

In ordinary gauge field theory, the most general 2-dimensional action on a multiplet of scalar fields is a nonlinear sigma model:

$$\int d^2x \quad g_{kl} \partial_a \phi^k \partial^a \phi^l$$

This has the interpretation of being a gauge field theory in which the field is a coordinate on a manifold M w. metric g_{kl}



The SUSY version above is a nonlinear sigma model on some sort of complex manifold.

A physicist's notion of a complex manifold is quite

particular. We image a manifold parametrized by complex
coordinates $z^k \bar{z}^k$

with Riemannian geometry and a Hermitian metric

$$ds^2 = g_{k\bar{l}} dz^k d\bar{z}^l$$

Mathematicians have a much more general notion of a complex
manifold. First, the manifold should have an even number of
dimensions ($2N$). Next, we need a way to sort coordinates into
holomorphic and anti-holomorphic. This can be done by postulating
a complex structure, a tensor J_i^j st.

$$(J^2)_i^k J_i^j J_j^k = -\delta_i^k$$

with equal numbers of $+i$ and $-i$ eigenvalues. If we
choose coordinates z, \bar{z} in the eigenspaces of J , then

$$J_i^j = +i \delta_i^j \quad J_{\bar{i}}^{\bar{j}} = -i \delta_{\bar{i}}^{\bar{j}} \quad i, j = 1-N$$

However, this refers only to one coordinate patch. To be a
complex structure, J must be defined globally (i.e. the patches
must fit together)

Now we have a complex manifold, but the metric could
still have the general form

$$ds^2 = g_{i\bar{j}} dz^i d\bar{z}^j + g_{\bar{i}j} d\bar{z}^i dz^j + g_{\bar{i}\bar{j}} d\bar{z}^i d\bar{z}^j$$

and the geometry might not be Riemannian. If we postulate

also a Hermitian metric and Γ_{km}^l a Riemannian connection for which J is covariantly constant, we obtain a Kähler manifold.

Given a Kähler manifold, let $J_{i\bar{k}} = J_i^j g_{j\bar{k}}$, and define the Kähler form

$$\begin{aligned} \omega &= J_{i\bar{k}} dz^i \wedge d\bar{z}^k \\ &= i g_{j\bar{k}} dz^j \wedge d\bar{z}^k \end{aligned}$$

if J is covariantly constant $\nabla_m J_i^j = 0$, with a symmetric connection, this implies,

$$\partial \omega = \bar{\partial} \omega = 0$$

with $\partial = dz^i \frac{\partial}{\partial z^i}$ $\bar{\partial} = d\bar{z}^{\bar{j}} \frac{\partial}{\partial \bar{z}^{\bar{j}}}$

so ω is a doubly closed form. Locally, then, ω can be integrated to

$$\omega = \partial \bar{\partial} K(z, \bar{z})$$

where K is a function called the Kähler potential.

Note that the "Kähler transformations"

$$K(z, \bar{z}) \rightarrow K(z, \bar{z}) + F'(z) + F(\bar{z})$$

yields the same Kähler form and ultimately, the same metric and geometry.

We now see that, for a Kähler manifold

$$g_{i\bar{j}} = K_{,i\bar{j}}$$

Let's also compute the connection and curvature:

$$\Gamma_{i\bar{k}}^{\bar{j}} = \frac{1}{2} g^{\bar{j}l} [\partial_i g_{kl} + \partial_k g_{il} - \partial_l g_{ik}] = 0$$

$$\Gamma_{\bar{i}k}^{\bar{j}} = \frac{1}{2} g^{\bar{j}l} [\partial_{\bar{i}} g_{kl} + \partial_k g_{\bar{i}l} - \partial_l g_{\bar{i}k}] = 0$$

with $g_{i\bar{l}} = K_{i\bar{l}}$ etc.

so i indices are not transformed to \bar{i} indices by parallel transport. ✓

$$\Gamma_{i\bar{k}}^{\bar{j}} = \frac{1}{2} g^{j\bar{l}} [\partial_i g_{k\bar{l}} + \partial_k g_{i\bar{l}} - \partial_{\bar{l}} g_{ik}]$$

$$= g^{j\bar{l}} K_{,ik\bar{l}}$$

$$\Gamma_{\bar{i}k}^{\bar{j}} = g^{\bar{j}l} K_{\bar{i}k\bar{l}}$$

the nonvanishing elements of the curvature tensor are:

$$R_{i\bar{j}}^{\bar{k}}{}_{\bar{l}} = \partial_i \Gamma_{\bar{j}}^{\bar{k}}{}_{\bar{l}} - \partial_{\bar{j}} \Gamma_i^{\bar{k}}{}_{\bar{l}} + \Gamma_i^{\bar{k}}{}_{\bar{m}} \Gamma_{\bar{j}}^{\bar{m}}{}_{\bar{l}} - \Gamma_{\bar{j}}^{\bar{k}}{}_{\bar{m}} \Gamma_i^{\bar{m}}{}_{\bar{l}}$$

$$= \partial_i \Gamma_{\bar{j}}^{\bar{k}}{}_{\bar{l}} = g^{\bar{k}m} K_{i\bar{j}m\bar{l}} - g^{\bar{k}m} \partial_i g_{m\bar{n}} g^{\bar{n}p} K_{\bar{j}\bar{l}p}$$

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$$R_{i\bar{j}k\bar{l}} = K_{i\bar{j}k\bar{l}} - K_{ik\bar{m}} g^{\bar{n}p} K_{\bar{j}\bar{l}p}$$

note that this expression has the symmetry of the Riemann

then $(i\bar{j}) \leftrightarrow (k\bar{l})$.

Now let's put this back into the Lagrangian on p.10. We now recognize that the $\partial_m \phi^{\dagger} \partial^m \phi$ piece is a nonlinear sigma model on a Kähler manifold with metric

$$g_{\bar{k}l} = K_{, \bar{k}l}$$

and Kähler potential $K(\phi, \phi^{\dagger})$. Note that the metric is not affected by a Kähler transformation

$$K(\phi, \phi^{\dagger}) \rightarrow K(\phi, \phi^{\dagger}) + F(\phi) + [F(\phi)]^*$$

In fact, the whole SUSY action is not affected, since $\int d^4\theta F(\Phi) = 0$ if $F(\Phi)$ is a chiral superfield. Using also

$$\Gamma_{lm}^n = g^{n\bar{k}} K_{, \bar{k}lm} \qquad g^{ab} = \text{inverse of } g_{ab}$$

The Lagrangian on p.10 becomes

$$\begin{aligned} & \int d^4\theta K(\Phi, \Phi^{\dagger}) \\ &= g_{\bar{k}l} \left\{ \partial_a \phi^{\dagger k} \partial^a \phi^l + \frac{i}{2} \psi^{k\dagger} \bar{\sigma}^a (\partial_a \psi^l + \Gamma_{mn}^l \partial_a \phi^m \psi^n) \right. \\ & \quad \left. - \frac{i}{2} (\partial_a \psi^{k\dagger} + \Gamma_{\bar{m}\bar{n}}^{\bar{k}} \psi^{\dagger m} \partial_a \phi^{\dagger n}) \bar{\sigma}^a \psi^l + F^{\dagger k} F^l \right\} \\ & \quad - \frac{1}{2} K_{\bar{k}lm} F^{\dagger k} \psi_c^{\dagger l} \psi_c^m + \frac{1}{2} K_{\bar{k}lm} \psi_c^{\dagger k} \psi_c^l F^m \\ & \quad - \frac{1}{4} K_{\bar{k}lmn} \psi_c^{\dagger k} \psi_c^{\dagger l} \psi_c^m \psi_c^n \end{aligned}$$

The derivative on the ψ field is

$$D_a \psi^l = \partial_a \psi^l + \Gamma_{mn}^l \partial_a \phi^m \psi^n$$

Interpreting ψ^l as a vector in the tangent space of M , this covariant derivative keeps track of the parallel transport of ψ relative to the coordinate directions $\partial_a \phi^m$. We can integrate D_a by parts:

$$\int d^4x \, g_{kl} (-D_a \psi^{k\dagger}) \psi^l = \int d^4x \, \psi^{k\dagger} D_a \psi^l$$

Finally, eliminating F , we find a 4-term of the structure

$$-\frac{1}{4} (K_{\bar{k}lmn} - K_{\bar{k}l\bar{p}} g^{\bar{p}\bar{q}} K_{\bar{q}mn}) \psi_c^{k\dagger} \psi^{l\dagger} \psi_c^{mT} \psi^n$$

which is just the Riemann curvature of M . So, finally, the most general kinetic ~~term~~ term for a chiral superfield is a ~~nonlinear sigma~~ model on a Kähler manifold

$$\mathcal{L} = \int d^4\theta \, \mathcal{K}(\Phi, \Phi^*)$$

$$= g_{kl} \left\{ \partial_a \phi^{*k} \partial^a \phi^l + \psi^{*k} i \bar{\sigma}^a \partial_a \psi^l \right\}$$

$$- \frac{1}{4} R_{\bar{i}jkl} \psi_c^{*i} \psi^{*j} \psi_c^{kT} \psi^l$$

In particular, if a supersymmetric model has a manifold of degenerate vacuum states with $N=1$ SUSY, those states form not only a complex manifold but a Kähler manifold.

④ Energy-momentum in $N=1$ SUSY

As a final application of superspace, I would like to explain what the equation of energy-momentum conservation looks like in $N=1$ SUSY.

Let's begin with an example, a model with a chiral superfield with superpotential $W(\Phi)$:

$$\int d^4\theta \Phi^* \Phi + \int d^3\theta W(\Phi) + h.c.$$

It is not so clear how to find the superspace equation of motion for Φ . Note that we cannot simply vary $\int d^4\theta \Phi^* \Phi$ freely with respect to Φ because Φ is a constrained field. However, we can rewrite the action as

$$\int d^2\theta \left\{ (-\frac{1}{4} \bar{D}^2) (\Phi^* \Phi) + W(\Phi) \right\} + \int d^2\bar{\theta} (W(\Phi))^*$$

and vary Φ freely under the integral $\int d^2\theta$. Notice that, in the first term, $\bar{D}^2(\Phi^* \Phi) = (\bar{D}^2 \Phi^*) \Phi$, since $\bar{D}_\alpha \Phi = 0$.

Then we can read off the equation of motion

$$-\frac{1}{4} \bar{D}^2 \Phi^* + \frac{\partial W}{\partial \Phi} = 0$$

Now let

$$J_{\alpha\beta} = +i(\bar{\sigma}\cdot\partial)_{\alpha\beta} \Phi^* \Phi - i\Phi^* (\bar{\sigma}\cdot\partial)_{\alpha\beta} \Phi + \frac{1}{2} \bar{D}_\alpha \Phi^* D_\beta \Phi$$

This object is reminiscent of the energy-momentum tensor in being a quadratic form in the components of the chiral field.

Using the equation of motion above, we can work out the equation of motion for $J_{\alpha\beta}$:

$$\begin{aligned} (\bar{D}J)_\beta &= -c_{\gamma\alpha} \bar{D}_\gamma J_{\alpha\beta} \\ &= -(c_{\gamma\alpha} \bar{D}_\gamma (\bar{\sigma}\cdot\partial)_{\alpha\beta} \Phi^*) \Phi + ic_{\gamma\alpha} \bar{D}_\gamma \Phi^* (\bar{\sigma}\cdot\partial)_{\alpha\beta} \Phi \\ &\quad - \frac{1}{2} (c_{\gamma\alpha} \bar{D}_\gamma \bar{D}_\alpha) \Phi^* D_\beta \Phi + \frac{1}{2} \bar{D}_\alpha \Phi^* c_{\gamma\alpha} \bar{D}_\gamma D_\beta \Phi \end{aligned}$$

using

$$\begin{aligned} \{\bar{D}_\alpha, D_\beta\} &= \frac{1}{4} [(-c_{\gamma\alpha} \bar{D}_\gamma \bar{D}_\alpha), D_\beta] \Phi^* \Phi + c_{\gamma\alpha} \bar{D}_\gamma \Phi^* \{\bar{D}_\alpha, D_\beta\} \Phi \\ &= 2i(\bar{\sigma}\cdot\partial)_{\alpha\beta} \Phi^* \Phi + c_{\gamma\alpha} \bar{D}_\gamma \Phi^* \{\bar{D}_\alpha, D_\beta\} \Phi \end{aligned}$$

The two terms with $\{\bar{D}_\alpha, D_\beta\} \Phi$ cancel (since $\bar{D}_\alpha \Phi = 0$) by the antisymmetry of $c_{\alpha\beta}$.

also $[\bar{D}^2, D_\beta] \Phi^* = -D_\beta \bar{D}^2 \Phi^*$

so

$$\begin{aligned} (\bar{D}J)_\beta &= 2 \frac{\partial W}{\partial \Phi} D_\beta \Phi - D_\beta \left(\frac{\partial W}{\partial \Phi} \right) \Phi \\ &= D_\beta \left(3W[\Phi] - \Phi \frac{\partial W}{\partial \Phi} \right) \end{aligned}$$

so, we find that

$$(\bar{D}J)_\beta = D_\beta S$$

where $J_{\alpha\beta}$ is a Hermitian, general superfield and

S is a chiral superfield. Ferrara and Zumino recognized that this is the general form of the equation of energy-momentum conservation in superspace. They provided several more examples, here is a simple one: The equation of motion of an Abelian gauge field is

$$-D\bar{D}^2 D V = 0 \quad \text{or} \quad D_\alpha W_\alpha = 0$$

This implies that

$$J_{\alpha\beta} = c W_\alpha^* (W_\beta)$$

$$\begin{aligned} \text{satisfies } (\bar{D}J)_\beta &= -c_{\gamma\alpha} \bar{D}_\gamma (c W_\alpha^*) (W_\beta) \\ &= -\bar{D}_\gamma W_\gamma^* (W_\beta) = 0 \end{aligned}$$

The Ferrara - Zumino equation is satisfied with the trivial RHS $S = 0$.

Actually, we can have $S = 0$ also in the case of the chiral superfield if the superpotential is

$$W(\Phi) = \lambda \Phi^3$$

This is the superpotential that gives a classically scale invariant theory, since λ is a dimensionless coeff constant, I would now like to show in some generality that

$$(\bar{D}J)_\beta = D_\beta S \Rightarrow \text{energy-momentum conservation}$$

$$S = 0 \Rightarrow \text{scale invariance.}$$

Introduce a component-field expansion of $J_{\alpha\beta}$ or of J_a defined by

$$J_{\alpha\beta} = \bar{\sigma}_{\alpha\beta}^a J_a$$

$$J_a = C_a(x) + \theta^T c \chi_a(x) + \theta^2 M_a(x) - \bar{\theta}^T c \chi_a^*(x) + \bar{\theta}^2 M_a^*(x)$$

$$+ \bar{\theta} \bar{\sigma}^b \theta v_{ab}(x) + \bar{\theta}^2 \theta^T \rho_a(x) - \theta^2 \bar{\theta}^T \rho_a^*(x)$$

$$+ \theta^2 \bar{\theta}^2 d_a(x)$$

Since we will differentiate with respect to $\bar{D} = -\frac{\partial}{\partial \bar{\theta}} + i\sigma \cdot \partial \theta$, it will be useful to rewrite this replacing χ by

$$y^a = \chi^a + i \bar{\theta} \sigma^a \theta \quad \text{s.t.} \quad \bar{D} y^a = 0$$

$$\begin{aligned}
J_a = & c_a(y) + \theta^T c \chi_a(y) + \theta^2 m_a(y) - \bar{\theta}^T c \chi_a^* + \bar{\theta}^2 m_a^* \\
& + \bar{\theta} \bar{\sigma}^b \theta (v_{ab} - i \partial_b c_a) \\
& + \bar{\theta}^2 \theta^T c (\rho_a + i/2 c (\bar{\sigma}^b)^T \partial_b \chi_a^*) - \theta^2 \bar{\theta}^T c (\rho_a^* + i/2 c \bar{\sigma}^b \partial_b \chi_a) \\
& + \theta^2 \bar{\theta}^2 (d_a - i/2 \partial_b v_{cb} - 1/4 \partial^2 c_a)
\end{aligned}$$

Then

$$\begin{aligned}
-\bar{D}_\gamma c_{\alpha\beta} J_{\alpha\beta} &= - (c \bar{\sigma}^a)_{\gamma\beta} \left(- \frac{\partial}{\partial \bar{\theta}^\gamma} \right) J_a(y) \\
&= (c \sigma^a)_{\beta\gamma} \left\{ (c \chi_a^*)_\gamma + 2(c \bar{\theta})_\gamma m_a^* - (\bar{\sigma}^b \theta)_\gamma (v_{ab} - i \partial_b c_a) \right. \\
&\quad + 2(c \bar{\theta})_\gamma \theta^T c (\rho_a + i/2 c (\bar{\sigma}^b)^T \partial_b \chi_a^*) \\
&\quad + \theta^2 [c (\rho_a^* + i/2 c \bar{\sigma}^b \partial_b \chi_a)]_\gamma \\
&\quad \left. + 2(c \bar{\theta})_\gamma \theta^2 (d_a - i/2 \partial_b v_{ab} - 1/4 \partial^2 c_a) \right\}
\end{aligned}$$

This should be compared with $D_\beta S$, with

$$S = s(y) + \theta^T c \psi(y) + \theta^2 f(y)$$

$$\text{Using } D_\beta \gamma^a = -2i (\bar{\theta} \bar{\sigma}^a)_\beta \quad ,$$

$$\begin{aligned}
D_\beta S = & (c \psi)_\beta + 2(c \theta)_\beta f - 2i (\bar{\theta} \bar{\sigma}^a)_\beta \partial_a S \\
& - 2i (\bar{\theta} \bar{\sigma}^a \partial_a)_\beta \theta^T c \psi - 2i (\bar{\theta} \bar{\sigma}^a \partial_a)_\beta \theta^2 f
\end{aligned}$$

Now we can compare terms in $(\bar{D} J)_\beta = D_\beta S$.

$$\underline{1}: -(\bar{\sigma}^a)^T \chi_a^* = c \psi \quad \rightarrow \quad \psi = c(\bar{\sigma}^a)^T \chi_a^*$$

$$\underline{\bar{\theta}}: -2((\bar{\sigma}^a)^T \bar{\theta})_\beta \mathcal{M}_a^* = -2i((\bar{\sigma}^a)^T \bar{\theta})_\beta \partial_a S$$

$$\rightarrow \quad \partial_a S = -i \mathcal{M}_a^*$$

$$\underline{\theta}: -(c \sigma^a \bar{\sigma}^b \theta)_\beta (v_{ab} - i \partial_b c_a) = 2(c \theta)_\beta f$$

$$\rightarrow \quad f = -\frac{1}{2} (v_{aa} - i \partial_a c_a)$$

We have now identified the components of S in terms of components of J .

Next, $D_\beta S$ has no $\underline{\theta}^2$ term; this implies

$$\rho_a^* = -i \frac{1}{2} c \bar{\sigma}^b \partial_b \chi_a$$

In the $\bar{\theta}\theta$ term, replace $\bar{\theta}_\alpha \theta_\beta = \frac{1}{2} (\sigma^c)_{\alpha\beta}^T \bar{\theta} \bar{\sigma}_c \theta$

$$2 c \sigma^a c \frac{1}{2} (\sigma^c)^T c (\rho_a + i \frac{1}{2} c (\bar{\sigma}^b)^T \partial_b \chi_a^*) (\bar{\theta} \bar{\sigma}_c \theta)$$

$$= -2i (\bar{\sigma}^a)^T (\frac{1}{2} \sigma^c)^T c \partial_a \psi (\bar{\theta} \bar{\sigma}_c \theta)$$

so

$$0 = \bar{\theta} \bar{\sigma}_c \theta \left\{ -(\bar{\sigma}^a)^T (\sigma^c)^T c \left(+\frac{1}{2} c (\bar{\sigma}^b)^T \partial_b \chi_a^* + \frac{i}{2} c (\bar{\sigma}^b)^T \partial_b \chi_a^* \right) + i (\bar{\sigma}^a)^T (\sigma^c)^T c c \bar{\sigma}^b \partial_a \chi_b^* \right\}$$

$$0 = \bar{\theta} \bar{\sigma}_c \theta \{ +i \} (\partial_b \chi_a^+ - \partial_a \chi_b^+) (\bar{\sigma}^b \sigma^c \bar{\sigma}^a)$$

$$\rightarrow \quad \bar{\sigma}^a \sigma^c \bar{\sigma}^b (\partial_b \chi_a - \partial_a \chi_b) = 0$$

multiply by (σ_c) :

$$\rightarrow (-2\sigma^a \bar{\sigma}^b) (\partial_b \chi_a - \partial_a \chi_b) = 0$$

$$(-2) \{ (2\eta^{ab} - \sigma^b \bar{\sigma}^a) \partial_b \chi_a - \sigma^a \bar{\sigma}^b \partial_a \chi_b \} = 0$$

$$\text{or } \partial_a \mathcal{J}^a = 0 \quad \text{where } \mathcal{J}^a = (\chi^a - \sigma^a \bar{\sigma}^b \chi_b)$$

Finally, the $\theta^2 \bar{\theta}$ term gives:

$$\begin{aligned} \theta^2 [-2(\bar{\sigma}^a)^T \bar{\theta}]_\beta (d_a - \frac{i}{2} \partial_b U_{ab} - \frac{1}{4} \partial^2 C_a) \\ = -2i\theta^2 [(\bar{\sigma}^a)^T \bar{\theta}]_\beta \partial_a f \end{aligned}$$

so

$$d_a - \frac{i}{2} \partial_b U_{ab} - \frac{1}{4} \partial^2 C_a + i \frac{1}{2} \partial_a (U_{bb} - i \partial_b C_b) = 0$$

The missing part is $\partial_b (v_{ab} - \eta_{ab} v_{cc}) = 0$

So if

$$t_{ab} = (v_{ab} - \eta_{ab} v_{cc})$$

$$\mathcal{J}_b = (\chi_b - \sigma_b \bar{\sigma} \cdot \chi)$$

the relation $(\bar{D}J)_\beta = D_\beta S$ implies $\partial_b \mathcal{J}^b = \partial_b t^{ab} = 0$

t_{ab} should be the energy-momentum tensor. The spin conserved current should be the supersymmetry current!

Now, what if $S = 0$? In that case,

$\psi = 0$ and $f = 0$. These relations imply

$$\bar{\sigma}^a J_a = 0 \quad t_a^a = 0 \quad \partial_a c^a = 0$$

The second of these relations is the equation for scale invariance.

If D^a is the dilatation current, $D^a = (t^{ab} x_b)$

$$\partial_a (D^a) = \partial_a (t^{ab} x_b) = t_a^a = 0$$

Scale invariance plus SUSY implies an additional fermionic symmetry called superconformal invariance, and this leads to

$\bar{\sigma}^a J_a = 0$. Finally, there is a new current that is conserved. This is the current of "R-symmetry", the U(1) symmetry on superfields:

$$\Phi(x, \theta) \rightarrow e^{i\frac{2}{3}\alpha} \Phi(x, e^{-i\alpha}\theta)$$

Notice that this is a symmetry of scale invariant superpotentials satisfying

$$\left(\frac{\partial W}{\partial \Phi_k}\right) \Phi_k = 3W$$

i.e. superpotentials cubic in superfields. Under R-symmetry

$$\phi_k \rightarrow e^{i\frac{2}{3}\alpha} \phi_k \quad \psi_k \rightarrow e^{-i\frac{1}{3}\alpha} \psi_k$$

The vector superfield transforms $V(x, \theta, \bar{\theta}) \rightarrow V(x, e^{-i\alpha}\theta, e^{+i\alpha}\bar{\theta})$

which implies

$$\lambda^a \rightarrow e^{+i\alpha} \lambda^a$$

Then, R symmetry is also a symmetry of the supersymmetric gauge interactions - eg.

$$\sqrt{2}g \phi^\dagger \lambda^{aT} \psi$$

Although Yang-Mills theory is classically scale invariant, this invariance is violated quantum-mechanically by explicit constant renormalization. This effect is described by the equation

$$t_a^a = \frac{1}{2g} \beta(g) (F_{mn}^a)^2$$

In SUSY, we recognize this equation as being part of a superfield relation that also implies nonconservation of the superconformal current of the R-current. I would like to claim that these results are summarized in the equation

$$S = \frac{\beta(g)}{g} \cdot \left(\frac{1}{2} W^T C W^a \right)$$

In particular, the Θ^2 component of this equation is

$$(U_{aa} - i \partial_a C^a) = \left(\frac{2\beta(g)}{3g} \right) \left[\frac{1}{4} (F_{mn}^a)^2 - \frac{2}{8} \epsilon^{mnpq} F_{mn}^a F_{pq}^a + \dots \right]$$

Now, if $J_{\alpha\beta} = (CW^\dagger)_\alpha (WC)_\beta$, we can show that

$$C^a = \frac{1}{2} \lambda^\dagger \bar{\sigma}^a \lambda = \frac{1}{2} J_R^a \quad \text{w. the normalization of}$$

$$\text{R-charges above, and} \quad U_{ab} \sim F^a C F^b + \dots$$

which is the usual nonadjoint of the adjoint tensor.

Since

$$t_{ab} = (V_{ab} - \eta_{ab} V^c_c), \quad -3V_{aa} = t_{aa}$$

Thus, the above equations also include (up to signs)

$$t_{aa} = \frac{\beta(g)}{2g} (F_{mn}^a)^2$$

as required.

In a non-SUSY gauge theory

$$\beta(g) = -\frac{g^3}{(4\pi)^2} \left[\frac{11}{3} C_2(G) - \sum_{\substack{\text{chiral} \\ \text{matter} \\ \text{fermions}}} \frac{2}{3} C(r_k) - \sum_{\substack{\text{complex} \\ \text{matter} \\ \text{scalars}}} \frac{1}{3} C(r_k) \right]$$

where

$$C(r) = \text{tr}[t^a t^b]_r$$

$$C_2(G) = C(\text{adjoint}) = N \text{ for } SU(N)$$

With SUSY, we should include the gauginos with the gauge fields:

$$\frac{11}{3} C_2(G) - \frac{2}{3} C_2(G) = 3 C_2(G)$$

and add the contributions of chiral fermions and scalars in a chiral supermultiplet. Then

$$\beta(g) = -\frac{g^3}{(4\pi)^2} \left[3 C_2(G) - \sum_{\substack{\text{chiral} \\ \text{supermult.}}} C(r_k) \right]$$

We now predict

$$\partial_a J_R^a = 2 \partial_a C^a = 2 \cdot \frac{\beta(g)}{12g} \epsilon^{mnpq} F_{mn}^a F_{pq}^a$$

or

$$\partial_a J_R^a = \frac{g^2}{32\pi^2} \left[C_2(G) - \sum_{\substack{\text{chiral} \\ \text{S.F.'s}}} \frac{1}{3} C(r_k) \right] \epsilon^{mnpq} F_{mn}^a F_{pq}^a$$

Now J_R is a chiral current, and so it should have a gauge anomaly!

$$\partial_a J_R^a = \frac{g^2}{32\pi^2} \text{tr} [R t^a t^b] \epsilon^{mnpq} F_{mn}^a F_{pq}^b$$

from $J_R = \sum_b \psi_b^\dagger \gamma_5 \psi_b + \dots$

The trace is over chiral fermion representations, with then R charges.

$$\text{tr} R t^a t^b = (d) C_2(G) \delta^{ab} + \sum_k (-\frac{1}{3}) C(r_k)$$

So the equation above is exactly right!

You might recall that, with a proper choice of regulator, the chiral anomaly is present in 1-loop diagrams only. The analogous regulator for a SUSY theory gives only a 1-loop contribution to the β function. However, the β function is conventionally defined with a different scheme, and so the conventional β function receives contributions of all orders. Relating the two schemes, Shifman, Vainshteyn, and Zakharov found the exact β -function of $N=1$ SUSY gauge theories. This result is nicely explained in Arkani-Hamed & Murayama, hep-th/9707133.