

11. Supergravity

In the previous lecture, I introduced the formulation of gravity in terms of vierbein and spin connection and the Rarita-Schwinger equation for a spin- $\frac{3}{2}$ field. Ferrara, Freedman, and van Nieuwenhuizen put the pieces together and showed that the resulting theory is locally supersymmetric. This is supergravity.

$$\mathcal{L} = \mathcal{L}_H + \mathcal{L}_{RS}$$

$$= -\frac{1}{2\kappa^2} e R - \frac{1}{2} \left[\psi_k^\dagger \epsilon^{klmn} \partial_l D_m \psi_n - D_m \psi_k^\dagger \epsilon^{klmn} \partial_l \psi_n \right]$$

viewed as a function of e_m^a , ω_m^{ab} , ψ_m^a . In this lecture, I will demonstrate that this theory has local $N=1$ SUSY.

To begin, let me determine ω_m^{ab} in this theory. As in the previous example, I will treat ω_m^{ab} as an independent field in \mathcal{L} and determine its equation by setting

$$\frac{\delta \mathcal{L}}{\delta \omega_m^{ab}} = 0.$$

This will be a constant equation that allows us to solve algebraically

for ω_m^{ab} , To study the variation w.r.t to ω , I would like to
 again rewrite the Hilbert action as

$$\begin{aligned} \mathcal{L}_H &= -\frac{1}{2\kappa^2} e^m{}_a e^n{}_b R_{mn}{}^{ab} \\ &= +\frac{1}{8\kappa^2} \epsilon^{klmn} \epsilon_{abcd} e^c{}_k e^d{}_l R_{mn}{}^{ab} \end{aligned}$$

then

$$\begin{aligned} \delta \mathcal{L} &= \frac{1}{8\kappa^2} \epsilon^{klmn} \epsilon_{abcd} e^c{}_k e^d{}_l D_m \delta \omega_n{}^{ab} \cdot 2 \\ &\quad - \frac{1}{2} \epsilon^{klmn} \left\{ \psi_k^+ \bar{\sigma}_l \frac{1}{2} \delta \omega_m{}^{ab} \sigma_{ab} \psi_n + \frac{1}{2} \delta \omega_m{}^{ab} \psi_k^+ \bar{\sigma}_{ab} \bar{\sigma}_l \psi_n \right\} \\ &= -\frac{1}{2\kappa^2} \epsilon^{klmn} \epsilon_{abcd} (D_m e^c{}_k) e^d{}_l \delta \omega_n{}^{ab} \\ &\quad - \frac{1}{4} \epsilon^{klmn} \delta \omega_m{}^{ab} e^d{}_l \psi_k^+ (\bar{\sigma}_d \sigma_{ab} + \bar{\sigma}_{ab} \bar{\sigma}_d) \psi_n \\ &= \frac{1}{2\kappa^2} \epsilon^{klmn} \delta \omega_m{}^{ab} e^d{}_l \left\{ \epsilon_{abcd} D_n e^c{}_k \right. \\ &\quad \left. - \frac{\kappa^2}{2} \psi_k^+ (\bar{\sigma}_d \sigma_{ab} + \bar{\sigma}_{ab} \bar{\sigma}_d) \psi_n \right\} \end{aligned}$$

now $\bar{\sigma}_{ab} \bar{\sigma}_d + \bar{\sigma}_d \sigma_{ab} = -i \epsilon_{abcd} \bar{\sigma}^c$

$$\begin{aligned} &= \frac{1}{2\kappa^2} \epsilon^{klmn} \epsilon_{abcd} \delta \omega_m{}^{ab} e^d{}_l \\ &\quad \cdot \left\{ -D_k e_n{}^c - i \frac{\kappa^2}{2} \psi_k^+ \bar{\sigma}^c \psi_n \right\} \end{aligned}$$

$$\text{so } T_{mn}^c = D_m e_n^c - D_n e_m^c = -i \frac{k^2}{2} (\psi_m^\dagger \bar{\sigma}^c \psi_n - \psi_n^\dagger \bar{\sigma}^c \psi_m) \quad 3$$

(more conveniently: $D_m e_n^c = -i \frac{k^2}{2} \psi_m^\dagger \bar{\sigma}^c \psi_n$)

The torsion is fixed to be nonzero and a natural bilinear in the gravitino fields. Nevertheless, we can solve explicitly for ω_m^{ab} in terms of e_m^a and ψ_m as indicated in the previous lecture.

Look back at the gravitino field equation from the previous lecture:

$$\epsilon^{klmn} \nabla_{ml}^b \bar{\sigma}_b \psi_n = -ik^2 \epsilon^{klmn} \bar{\sigma}_b \psi_n \psi_m^\dagger \bar{\sigma}^b \psi_l$$

$$\text{Fierz id.} = -ik^2 \epsilon^{klmn} (-2) \psi_m^\dagger (\psi_l^T \psi_n)$$

$$= 0$$

since $\psi_l^T \psi_n$ is symmetric under $(l \leftrightarrow n)$

so the gravitino field equation is

$$-\epsilon^{klmn} \bar{\sigma}_l D_m \psi_n = 0$$

as promised.

Now I would like to show that \mathcal{L} is supersymmetric.

In fact, we will consider local supersymmetry, with a spin parameter $\xi(x)$ that depends on x . The transform

laws are:

$$\delta_{\xi} e_m^a = \kappa \left(\xi^{\dagger} \bar{\sigma}^a \psi_m + \psi_m^{\dagger} \bar{\sigma}^a \xi \right)$$

$$\delta_{\xi} \psi_m = -2 \frac{i}{\kappa} D_m \xi$$

The transformation of e_m^a is similar to that of λ_m in an $N=1$ SUSY sys. However, now ξ is a function of x . The transformation of ψ_m includes dependence on $D_m \xi$, so ψ_m is the gauged field of local SUSY. Notice that the SUSY invariance will generate the gauge invariance of the Rarita-Schwinger equation.

What is the transformation law for ω_m^{ab} ? This does not have to be postulated independently. Rather, we solve for ω_m^{ab} in terms of e_m^a and ψ_m using the torsion constraint and then plug in the above expressions. After a lengthy computation, this gives

$$\delta_{\xi} \omega_m^{bc} = \frac{\kappa}{2} e_m^a \left\{ \xi^{\dagger} \bar{\sigma}_c \psi_{ab} + \xi^{\dagger} \bar{\sigma}_b \psi_{ca} - \xi^{\dagger} \bar{\sigma}_a \psi_{bc} + \psi_{ab}^{\dagger} \bar{\sigma}_c \xi + \psi_{ca}^{\dagger} \bar{\sigma}_b \xi - \psi_{bc}^{\dagger} \bar{\sigma}_a \xi \right\}$$

where $\psi_{ab} = e_a^m e_b^n (D_m \psi_n - D_n \psi_m)$. The minus signs \uparrow look odd, but they are correct.

Now I would like to show that

$$(1) \quad [\delta_\xi, \delta_\eta] = 2i (\xi^\dagger \bar{\sigma}^m \eta - \eta^\dagger \bar{\sigma}^m \xi) \partial_m + (\text{possible gauge terms})$$

$$(2) \quad \delta_\xi \mathcal{L} = 0$$

As before, (2) is easier and I'll do that first.

In computing $\delta_\xi \mathcal{L}$, we have three terms:

$$\delta_\xi \mathcal{L} = \frac{\delta \mathcal{L}}{\delta e_m^a} \delta_\xi e_m^a + \frac{\delta \mathcal{L}}{\delta \omega_m^{ab}} \delta_\xi \omega_m^{ab} + \frac{\delta \mathcal{L}}{\delta \psi_m} \delta_\xi \psi_m$$

However, ω_m^{ab} has been determined by the condition $\frac{\delta \mathcal{L}}{\delta \omega_m^{ab}} = 0$, so the second term vanishes. We need only compute the terms w. explicit variation of e_m^a or ψ_m .

Then

$$\delta_\xi \mathcal{L}_H = \frac{1}{4k^2} \epsilon^{klmn} \epsilon_{abcd} \delta e_k^c e_l^d R_{mn}^{ab}$$

$$= \frac{1}{4k^2} \epsilon^{klmn} \epsilon_{abcd} \cdot k \cdot (\xi^\dagger \bar{\sigma}^c \psi_k + \psi_k^\dagger \bar{\sigma}^c \xi) e_l^d R_{mn}^{ab}$$

The variation of \mathcal{L}_H is somewhat more complicated. We have to vary ψ_k^\dagger, ψ_n , and also the e_l^d inside $\bar{\sigma}_l = e_l^d \bar{\sigma}_d$. I would like to do this in two steps: first ignore the presence of this e_l^d and then correct for it. Call these pieces (a) and (b)

$$\begin{aligned} \delta_\xi \mathcal{L}_{PS}(a) = & -\frac{1}{2} \left[\psi_k^+ \epsilon^{klmn} \bar{\psi}_l D_m \left(-\frac{2i}{k} D_n \xi \right) \right. \\ & + \left(\frac{2i}{k} D_k \xi^+ \right) \epsilon^{klmn} \bar{\psi}_l D_m \psi_n \\ & - D_m \psi_k^+ \epsilon^{klmn} \bar{\psi}_l \left(-\frac{2i}{k} D_n \xi \right) \\ & \left. - \left(\frac{2i}{k} D_m D_k \xi^+ \right) \epsilon^{klmn} \bar{\psi}_l \psi_n \right] \end{aligned}$$

It will be useful to integrate by parts in the middle two terms. To do this, we have to pass D_k or D_n through e_l^d . Put these terms into $\delta_\xi \mathcal{L}_{PS}(b)$. With this revision

$$\begin{aligned} \delta_\xi \mathcal{L}_{PS}(a) = & \left(-\frac{1}{2} \right) \left(-\frac{2i}{k} \right) \epsilon^{klmn} \\ & \left\{ \psi_k^+ \bar{\psi}_l D_m D_n \xi + \xi^+ \bar{\psi}_l D_k D_m \psi_n \right. \\ & \left. + (D_n D_m \psi_k^+) \bar{\psi}_l \xi + D_m D_k \xi^+ \bar{\psi}_l \psi_n \right\} \end{aligned}$$

now replace

$$\begin{aligned} [D_m D_n] \xi &= \frac{1}{2} [D_m, D_n] \xi = \frac{1}{2} \cdot \frac{1}{2} R_{mn}^{ab} \sigma_{ab} \xi \\ &= \frac{i}{k} \epsilon^{klmn} e_l^d \left\{ \psi_k^+ \bar{\psi}_d \frac{1}{4} R_{mn}^{ab} \sigma_{ab} \xi + \xi^+ \bar{\psi}_d \frac{1}{4} R_{km}^{ab} \sigma_{ab} \psi_n \right. \\ & \quad \left. - \frac{1}{4} R_{nm}^{ab} \psi_k^+ \bar{\psi}_{ab} \bar{\psi}_d \xi - \frac{1}{4} R_{mk}^{ab} \xi^+ \bar{\psi}_{ab} \bar{\psi}_d \psi_n \right\} \end{aligned}$$

$$\delta_\xi \mathcal{L}_{RS}(a) = \frac{i}{4K} \epsilon^{klmn} e_l^d R_{mn}^{ab}$$

$$\left\{ \psi_k^+ \bar{\sigma}_d \sigma_{ab} \xi + \psi_k^+ \bar{\sigma}_{ab} \bar{\sigma}_d \xi \right. \\ \left. + \xi^+ \bar{\sigma}_d \sigma_{ab} \psi_k + \xi^+ \bar{\sigma}_{ab} \bar{\sigma}_d \psi_k \right\}$$

now use $\bar{\sigma}_{ab} \bar{\sigma}_d + \bar{\sigma}_d \sigma_{ab} = -i \epsilon_{abcd} \bar{\sigma}^c$
 $= +i \epsilon_{abcd} \bar{\sigma}^c$

$$= -\frac{1}{4K} \epsilon^{klmn} e_l^d R_{mn}^{ab} \epsilon_{abcd}$$

$$\left\{ \psi_k^+ \bar{\sigma}^c \xi + \xi^+ \bar{\sigma}^c \psi_k \right\}$$

so this precisely cancels $\delta_\xi \mathcal{L}_H$. Now we just have to clean up

$$\delta_\xi \mathcal{L}_{RS}(b) = (\text{terms from integrate by parts}) \\ + (\text{terms from } \delta e_l^d \text{ in } \mathcal{L}_{RS})$$

$$= \left(-\frac{1}{2} \right) \left(-\frac{2i}{K} \right) \epsilon^{klmnt} \frac{1}{3} (D_k e_l^d) \bar{\sigma}_d D_m \psi_n \\ + \left(-\frac{1}{2} \right) \left(-\frac{2i}{K} \right) \epsilon^{klmn} (D_m \psi_k^+) (D_n e_l^d) \bar{\sigma}_d \xi \right] \text{int. by parts.}$$

$$- \frac{1}{2} \epsilon^{klmn} \left(\psi_k^+ \bar{\sigma}_d D_m \psi_n - D_m \psi_k^+ \bar{\sigma}_d \psi_n \right) K \left(\xi^+ \bar{\sigma}_d \psi_l + \psi_l^+ \bar{\sigma}_d \xi \right)$$

$$\begin{aligned} \delta_S \mathcal{L}_{RS}^{(b)} &= \frac{i}{k} \epsilon^{klmn} (\xi^\dagger \bar{\sigma}_d D_m \psi_n) (D_k e_l^d) \\ &+ \frac{i}{k} \epsilon^{klmn} (D_m \psi_k^\dagger \bar{\sigma}_d \xi) (D_n e_l^d) \\ &- \frac{k}{2} \epsilon^{klmn} (\psi_k^\dagger \bar{\sigma}_d D_m \psi_n - D_m \psi_k^\dagger \bar{\sigma}_d \psi_n) (\xi^\dagger \bar{\sigma}^d \psi_l + \psi_l^\dagger \bar{\sigma}^d \xi) \end{aligned}$$

In the first two terms, replace

$$D_k e_l^d = -i \frac{k^2}{2} \psi_k^\dagger \bar{\sigma}^c \psi_l$$

In the last term, Fierz! Note that

$$\begin{aligned} \psi_k^\dagger \bar{\sigma}_d D_m \psi_n \quad \psi_l^\dagger \bar{\sigma}^d \xi &= -2 \psi_{kc}^\dagger \psi_l^* \xi^T_c D_m \psi_n \\ &= 0 \end{aligned}$$

since $\psi_{kc}^\dagger \psi_l^*$ is symmetric under $(k \leftrightarrow l)$

similarly, the term with $\bar{\sigma}_d \psi_n \dots \bar{\sigma}^d \psi_l$ vanishes. What remains is

$$\begin{aligned} \delta_S \mathcal{L}_{RS}^{(b)} &= \frac{i}{k} \epsilon^{klmn} (\xi^\dagger \bar{\sigma}_d D_m \psi_n) \left(-i \frac{k^2}{2}\right) (\psi_k^\dagger \bar{\sigma}^d \psi_l) \\ &+ \frac{i}{k} \epsilon^{klmn} (D_m \psi_n^\dagger \bar{\sigma}_d \xi) \left(-i \frac{k^2}{2}\right) (\psi_n^\dagger \bar{\sigma}^d \psi_l) \\ &- \frac{k}{2} \epsilon^{klmn} \left\{ \psi_k^\dagger \bar{\sigma}^d \psi_l \xi^\dagger \bar{\sigma}_d D_m \psi_n - D_m \psi_n^\dagger \bar{\sigma}_d \xi \psi_l^\dagger \bar{\sigma}^d \psi_n \right\} \\ &= \frac{k}{2} \epsilon^{klmn} \left\{ (\xi^\dagger \bar{\sigma}_d D_m \psi_n) (\psi_k^\dagger \bar{\sigma}^d \psi_l) - (\xi^\dagger \bar{\sigma}_d D_m \psi_n) (\psi_k^\dagger \bar{\sigma}^d \psi_l) \right. \\ &\quad \left. + (D_m \psi_n^\dagger \bar{\sigma}_d \xi) (\psi_n^\dagger \bar{\sigma}^d \psi_l) + (D_m \psi_n^\dagger \bar{\sigma}_d \xi) (\psi_l^\dagger \bar{\sigma}^d \psi_n) \right\} \end{aligned}$$

and everything cancels! Thus, finally

$$\delta_{\xi} \mathcal{L} = 0$$

Now study the closure of the SUSY commutators. Let's compute

$$\begin{aligned} [\delta_{\xi}, \delta_{\eta}] e_m^a &= \delta_{\xi} \cdot \kappa (\eta^{\dagger} \bar{\sigma}^a \psi_m + \psi_m^{\dagger} \bar{\sigma}^a \eta) - (\xi \leftrightarrow \eta) \\ &= \kappa \left\{ \eta^{\dagger} \bar{\sigma}^a \left(-\frac{2i}{\kappa} \right) D_m \xi + \frac{2i}{\kappa} D_m \xi^{\dagger} \bar{\sigma}^a \eta \right\} - (\xi \leftrightarrow \eta) \\ &= 2i \left[(D_m \xi^{\dagger}) \bar{\sigma}^a \eta + \xi^{\dagger} \bar{\sigma}^a (D_m \eta) \right] - (\xi \leftrightarrow \eta) \end{aligned}$$

I would like to manipulate this expression in a somewhat non-obvious

way.

$$\begin{aligned} &= 2i \partial_m (\xi^{\dagger} \bar{\sigma}^a \eta) + 2i \frac{1}{2} \omega_m^{cd} \xi^{\dagger} (\bar{\sigma}^a \sigma_{cd} - \bar{\sigma}_{cd} \bar{\sigma}^a) \eta \\ &\quad - (\xi \leftrightarrow \eta) \end{aligned}$$

now $\bar{\sigma}^a \sigma_{cd} - \bar{\sigma}_{cd} \bar{\sigma}^a = \eta_{ac} \bar{\sigma}_d - \eta_{ad} \bar{\sigma}_c$

$$= 2i \partial_m (\xi^{\dagger} \bar{\sigma}^n \eta e_n^a) + 2i \omega_m^a b \xi^{\dagger} \bar{\sigma}^b \eta - (\xi \leftrightarrow \eta)$$

$$\begin{aligned} &= 2i \partial_m (\xi^{\dagger} \bar{\sigma}^n \eta) e_n^a + 2i \xi^{\dagger} \bar{\sigma}^n \eta (\partial_m e_n^a) \\ &\quad + 2i \omega_m^a b \xi^{\dagger} \bar{\sigma}^b \eta - (\xi \leftrightarrow \eta) \end{aligned}$$

$$\begin{aligned}
&= \partial_m (2i \xi^{\dagger} \bar{\sigma}^n \eta) e_n^a + 2i (\xi^{\dagger} \bar{\sigma}^n \eta) \partial_n e_m^a \\
&\quad + 2i (\xi^{\dagger} \bar{\sigma}^n \eta) (\partial_m e_n^a - \partial_n e_m^a) \\
&\quad + 2i (\xi^{\dagger} \bar{\sigma}^n \eta) \omega_m^{ab} e_n^b - 2i \xi^{\dagger} \bar{\sigma}^n \eta \omega_n^{ab} e_m^b \\
&\quad + 2i \xi^{\dagger} \bar{\sigma}^n \eta \omega_n^{ab} e_m^b \\
&\quad - (\xi \leftrightarrow \eta)
\end{aligned}$$

m = term added
+ subtracted

$$\begin{aligned}
&= \partial_m (2i \xi^{\dagger} \bar{\sigma}^n \eta) e_n^a + (2i \xi^{\dagger} \bar{\sigma}^n \eta) \partial_n e_m^a \\
&\quad + 2i (\xi^{\dagger} \bar{\sigma}^n \eta) (D_m e_n^a - D_n e_m^a) \\
&\quad + 2i \xi^{\dagger} \bar{\sigma}^n \eta \omega_n^{ab} e_m^b - (\xi \leftrightarrow \eta)
\end{aligned}$$

• substitute for the torsion

$$\begin{aligned}
&= \partial_m (2i \xi^{\dagger} \bar{\sigma}^n \eta) e_n^a + (2i \xi^{\dagger} \bar{\sigma}^n \eta) \partial_n e_m^a \\
&\quad + 2i (\xi^{\dagger} \bar{\sigma}^n \eta) \left(-\frac{i k^2}{2}\right) (\psi_m^{\dagger} \bar{\sigma}^a \psi_n - \psi_n^{\dagger} \bar{\sigma}^a \psi_m) \\
&\quad + (2i \xi^{\dagger} \bar{\sigma}^n \eta) \omega_n^{ab} e_{mb} - (\xi \leftrightarrow \eta)
\end{aligned}$$

this expression has now come into the following form:

$$[\delta_\xi, \delta_\eta] e_m^a = \partial_m(a^n) e_n^a + a^n \partial_n e_m^a + \kappa (\Xi^+ \bar{\sigma}^a \psi_m + \psi_m^+ \bar{\sigma}^a \Xi) + \Omega^{ab} e_{mb}$$

The first line is a local translation with parameter

$$a^n(x) = 2i (\xi^+ \bar{\sigma}^n \eta - \eta^+ \bar{\sigma}^n \xi) \quad \text{as required}$$

The second line is a local SUSY transformation with parameter

$$\Xi(x) = \kappa (\xi^+ \bar{\sigma}^n \eta - \eta^+ \bar{\sigma}^n \xi) \psi_n$$

similar to the gauge term is $[\delta_\xi, \delta_\eta] A_m$ in the SUSY acts on a gauge field. The third term is a local Lorentz rotation with parameter

$$\Omega^{ab}(x) = 2i (\xi^+ \bar{\sigma}^n \eta - \eta^+ \bar{\sigma}^n \xi) \omega_n^{ab}$$

so the commutator of local SUSY transformations closes onto a combination of local translations, local SUSY, and local Lorentz, at least on e_m^a . To complete the proof of closure, we need to show:

$$[\delta_\xi, \delta_\eta] \psi_m = \partial_m(a^n) \psi_n + a^n \partial_n \psi_m - \frac{2i}{\kappa} D_m \Xi + \frac{1}{2} \Omega^{ab} \sigma_{ab} \psi_m$$

with the same parameters as above.

In fact, we will not be able to show this without using the gravitino equations of motion. In the system as we have described it, the SUSY commutators close only on shell. However, assume the gravitino equations proceed.

$$\begin{aligned}
[\delta_\xi, \delta_\eta] \psi_m &= \delta_\xi \left(-\frac{2i}{K} D_m \eta \right) - (\xi \leftrightarrow \eta) \\
&= \left(-\frac{2i}{K} \right) \frac{1}{2} (\delta_\xi \omega_m^{bc}) \sigma_{bc} \eta - (\xi \leftrightarrow \eta) \\
&= -\frac{i}{K} (\sigma_{bc} \eta) (\delta_\xi \omega_m^{bc}) - (\xi \leftrightarrow \eta)
\end{aligned}$$

now, I gave $\delta_\xi \omega_m^{bc}$ on p. 4. To make progress, it will be useful to simplify that expression using the field equation in the form

$$\delta_c \psi_{ab} + \delta_b \psi_{ca} + \delta_a \psi_{bc} = 0$$

so, with the field equation

$$\begin{aligned}
\delta_\xi \omega_m^{bc} &= \chi e_m^a \left\{ \xi^+ \delta_c \psi_{ab} + \xi^+ \delta_b \psi_{ca} \right. \\
&\quad \left. + \psi_{ab}^+ \delta_c \xi + \psi_{ca}^+ \delta_b \xi \right\}
\end{aligned}$$

since this is contracted with a structure antisymmetric

made $(b \leftrightarrow c)$

$$[\delta_\xi, \delta_\eta] \Psi_m = -2i (\sigma_{bc} \eta) e_m^a \left\{ \xi^\dagger \bar{\sigma}_c \psi_{ab} + \psi_{ab}^\dagger \bar{\sigma}_c \xi \right\} - (\xi \leftrightarrow \eta)$$

next rewrite

$$\begin{aligned} \sigma_{bc} \eta &= \frac{1}{4} [\sigma_b \bar{\sigma}_c - \sigma_c \bar{\sigma}_b] \eta \\ &= \frac{1}{2} \sigma_b \bar{\sigma}_c \eta - \frac{1}{2} \eta_{bc} \eta \end{aligned}$$

and use the field equation again in the form $\bar{\sigma}^c \psi_{ac} = 0$

$$= -i \sigma_b \bar{\sigma}_c \eta (\xi^\dagger \bar{\sigma}_c \psi_{ab} + \psi_{ab}^\dagger \bar{\sigma}_c \xi) e_m^a - (\xi \leftrightarrow \eta)$$

now Fierz; the term with

$$\bar{\sigma}_c \eta \psi_{ab}^\dagger \bar{\sigma}_c \xi = -2 \epsilon \psi_{ab}^{\dagger c} \xi^T c \eta$$

$\rightarrow 0$ after antisymmetry on $(\xi \leftrightarrow \eta)$

the other term gives

$$= -i (\xi^\dagger \bar{\sigma}_c \eta) (\sigma^b \bar{\sigma}_c \psi_{ab}) e_m^a - (\xi \leftrightarrow \eta)$$

Finally, replace

$$\sigma^b \bar{\sigma}_c \psi_{ab} = 2 \eta_c^b \psi_{ab} - \sigma_c^{\dagger b} \bar{\sigma}^b \psi_{ab}$$

and use the field equation to set $\bar{\sigma}^b \psi_{ab} = 0$.

Then

$$\begin{aligned}
[\delta_\xi, \delta_\eta] \psi_m &= -2i (\xi^\dagger \bar{\sigma}^c \eta - \eta^\dagger \bar{\sigma}^c \xi) e_m^a \psi_{ac} \\
&= -2i (\xi^\dagger \bar{\sigma}^n \eta - \eta^\dagger \bar{\sigma}^n \xi) \psi_{mn} \\
&= 2i (\xi^\dagger \bar{\sigma}^n \eta - \eta^\dagger \bar{\sigma}^n \xi) (\mathcal{D}_n \psi_m - \mathcal{D}_m \psi_n) \\
&= 2i (\xi^\dagger \bar{\sigma}^n \eta - \eta^\dagger \bar{\sigma}^n \xi) \partial_n \psi_m \\
&\quad + 2i (\xi^\dagger \bar{\sigma}^n \eta - \eta^\dagger \bar{\sigma}^n \xi) \frac{1}{2} \omega_n^{ab} \sigma_{ab} \psi_m \\
&\quad - 2i \mathcal{D}_m ((\xi^\dagger \bar{\sigma}^n \eta - \eta^\dagger \bar{\sigma}^n \xi) \psi_n) \\
&\quad + 2i \partial_m (\xi^\dagger \bar{\sigma}^n \eta - \eta^\dagger \bar{\sigma}^n \xi) \psi_n
\end{aligned}$$

this has now taken the correct form.

$$\begin{aligned}
[\delta_\xi, \delta_\eta] \psi_m &= a^n \partial_n \psi_m + \frac{1}{2} \Omega^{ab} \sigma_{ab} \psi_m \\
&\quad - \frac{2i}{k} \mathcal{D}_m \square + \partial_m (a^n) \psi_n
\end{aligned}$$

and our proof is complete.

To couple supergravity to matter, it will be very useful to have a system that realizes local SUSY off-shell. For this, we must add some auxiliary fields to the physical supergravity fields e_m^a, ψ_m . It is not so hard to count the off-shell degrees of freedom needed:

$$e_m^a = 16 \text{ components} - (6 + 4 \text{ gage degrees of freedom}) = 6$$

local Lorentz
local translation

$$\psi_m = 4 \times 2 \text{ complex} = 16 \text{ real components}$$

$$- (2 \text{ complex gage d.o.f freedom}) = 12$$

= 4 real \uparrow
 Rarita-Schwinger

So we need 6 more bosonic auxiliary fields. A possible choice is a vector plus a complex scalar:

$$e_m^a \quad \psi_m \quad b_m \quad M$$

I will describe the theory of this system in a future lecture.